

PHYSICS BEYOND THE SM

UNDER THE HIGGS LAMPPOST

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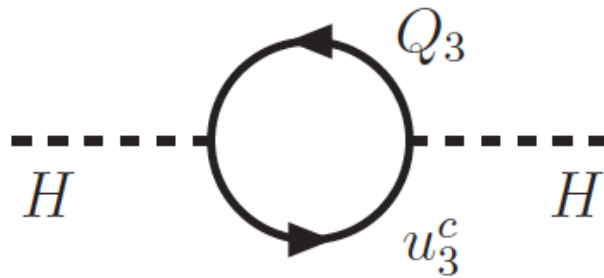


PHYSICS BEYOND THE SM

UNDER THE HIGGS LAMPPOST

1. The Quest for the SM & Beyond
2. A Strongly-coupled EW Sector
3. A Weakly-coupled Extension
4. Flavors of Matter Fields & EFT

The BIG Hierarchy Problem



Why the weak scale M_W is so much smaller than M_{Pl} ?

1. Dynamical generation: dimensional transmutation
2. Controlled by symmetry:
 - Shift symmetry (Nambu-Goldstone boson)
 - Super-symmetry (elementary particles)
 - Discrete symmetries

Supersymmetry: an elegant formalism

Julius Wess & Bruno Zumino (in 1974)

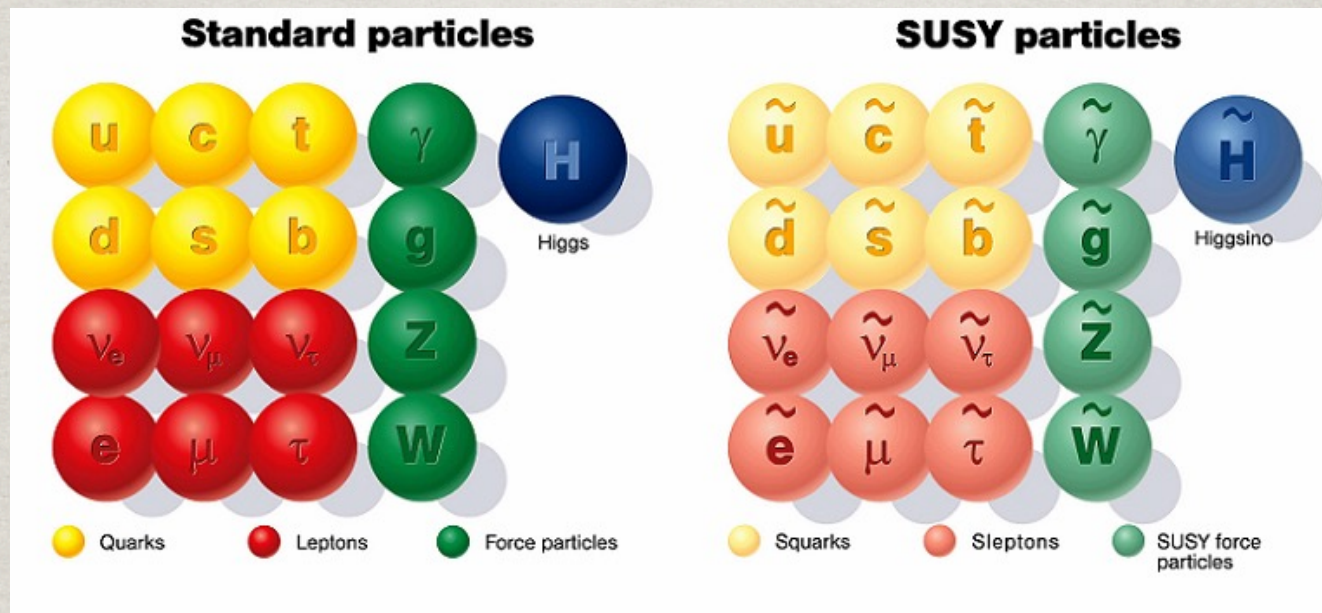
- Generalized extension of Poincare group of spacetime symmetry
- Critical ingredients for string theory:
World sheet/Spacetime SUSY
- Low energy effective field theory from compactification

- Eliminate the cosmological constant upto symmetry breaking
- Stabilizes weak scale at all orders
- Potential Grand Unification of EW & QCD
- Natural WIMP dark matter candidate

Supersymmetry: fermion \leftrightarrow boson

The SUSY generators Q transform fermions into bosons and vice-versa

$$Q|\text{Fermion}\rangle = |\text{Boson}\rangle, \quad Q|\text{Boson}\rangle = |\text{Fermion}\rangle$$



Under SUSY, the two partners must

- Have the same mass, the same coupling
- Spin different by $\frac{1}{2}$

Since we have not seen a partner of same mass, SUSY must be broken, at a higher scale.

Minimal Supersymmetric Standard-Model (MSSM)

Extended symmetry between *opposite* spin & statistics

particles	symbol	spin	mass param.
gluino	\tilde{g}	1/2	M_3
charginos	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^\pm$	1/2	M_2
neutralinos	$\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$	1/2	M_1, μ, B $m_{H_u}^2, m_{H_d}^2$
sleptons	$\tilde{e}_L, \tilde{\nu}_{eL}, \tilde{e}_R$	0	$m_{\ell L}^2$
	$\tilde{\mu}_L, \tilde{\nu}_{\mu L}, \tilde{\mu}_R$	0	
	$\tilde{\tau}_1, \tilde{\tau}_2, \tilde{\nu}_{\tau L}$	0	$m_{\ell R}^2$
squarks	$\tilde{u}_L, \tilde{d}_L, \tilde{u}_R, \tilde{d}_R$	0	m_{qL}^2
	$\tilde{c}_L, \tilde{s}_L, \tilde{c}_R, \tilde{s}_R$	0	
	$\tilde{t}_1, \tilde{t}_2, \tilde{b}_1, \tilde{b}_2$	0	m_{qR}^2
Higgs	h^0, H^0, A^0, H^\pm	0	$m_A^2, \tan \beta$

\tilde{t} versus t ; \tilde{W} versus W ; \tilde{H} versus H ; H_d versus H_u ,

$$\Delta m_H^2 \sim (M_{SUSY}^2 - M_{SM}^2) \frac{\lambda_f^2}{16\pi^2} \ln \left(\frac{\Lambda}{M_{SUSY}} \right).$$

only if the “soft-SUSY breaking”: $M_{SUSY} \sim \mathcal{O}(M_{SM})$.

→ “Weak scale” supersymmetry!

What about M_{SUSY} (in a hidden sector)?

- × Supersymmetry breaking mechanism is unknown.

Fermionic masses:

$$M_1, M_2, M_3, \mu \rightarrow M_{\chi_{1,2}^\pm}, M_{\chi_{1,2,3,4}^0};$$

Scalar masses:

$$M_{\tilde{q}_{L,R}}, M_{\tilde{l}_{L,R}};$$

Mixings:

$$\tan \beta, \sin \alpha \dots \dots$$

CP Phases:

$$\phi_{1,2,3\mu} \dots \dots$$

Parameter count in the SM and MSSM (no m'_ν s yet)

model	masses and mixing ang.	CP-viol. phases	TOTAL
SM	17	2	19
MSSM	79	45	124
(MSSM) _{BV}	97	62	159
(MSSM) _{LV}	157	122	279
(MSSM) _{BLV}	175	140	315

Based on observation:

* Proton stability:

⇒ R -parity conservation; or B, L not broken simultaneously (in 1st, 2nd generations).

* No excessively large CP-violation/FCNC:

⇒ no (or small) phases; sfermion mass degenerate (or heavy).

General parameter selection: Phenomenological MSSM (pMSSM)

The 20 dimensional pMSSM parameter space then includes

$$\begin{aligned} &M_1, M_2, M_3, \\ &m_{Q_1}, m_{U_1}, m_{D_1}, m_{L_1}, m_{E_1}, \\ &m_{Q_3}, m_{U_3}, m_{D_3}, m_{L_3}, m_{E_3}, \\ &A_t, A_b, A_\tau, \\ &m_{H_u}^2, m_{H_d}^2, \mu, B. \end{aligned}$$

scan over parameters



University of California, Santa Cruz

August 18-23, 2025

SUSY *is* still relevant!

Hitoshi Murayama (Berkeley, Kavli IPMU)

SUSY 2025, Santa Cruz, Aug 23, 2025

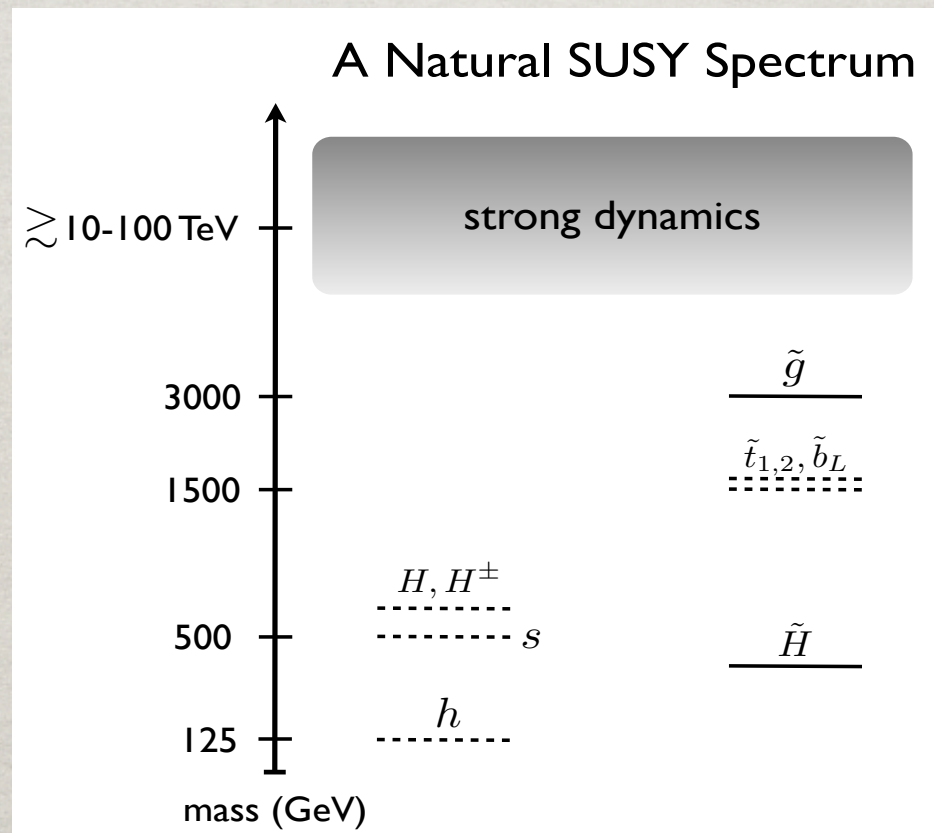
“Natural SUSY”: Radiative EWSB

✓ radiative EWSB by the large top Yukawa coupling:

$$M_Z^2/2 = \frac{m_{H_d}^2 - m_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} - \mu^2.$$

Assume higher scales correlated,
only insist on light Higgsinos:

Cohen, Kaplan, Nelson, 1996
Hall, Pinner, Ruderman, 2012
Baer, Barger, Huang, Tata, 2012



Higgs mass corrections & Heavy Higgs bosons

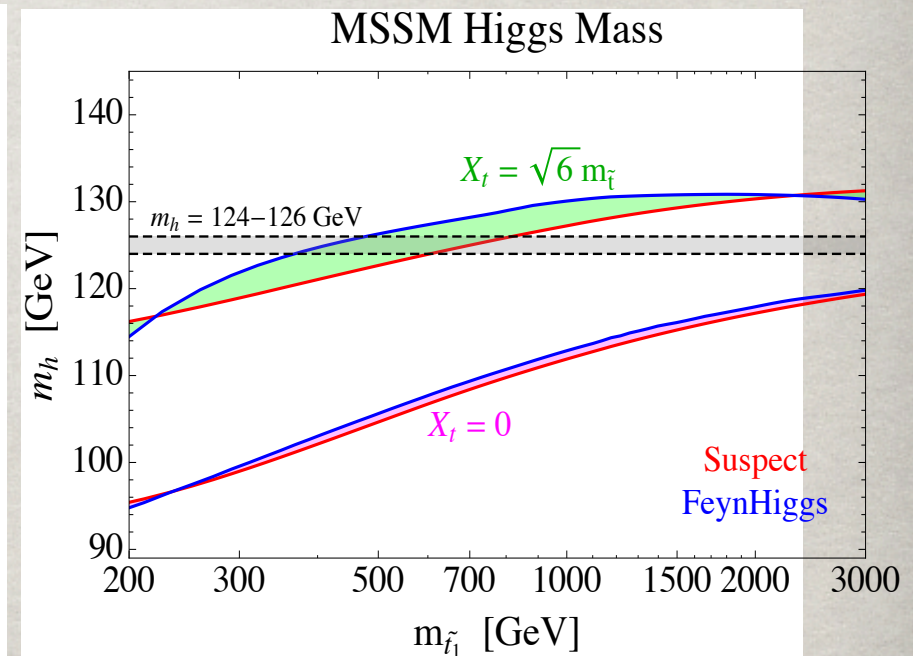
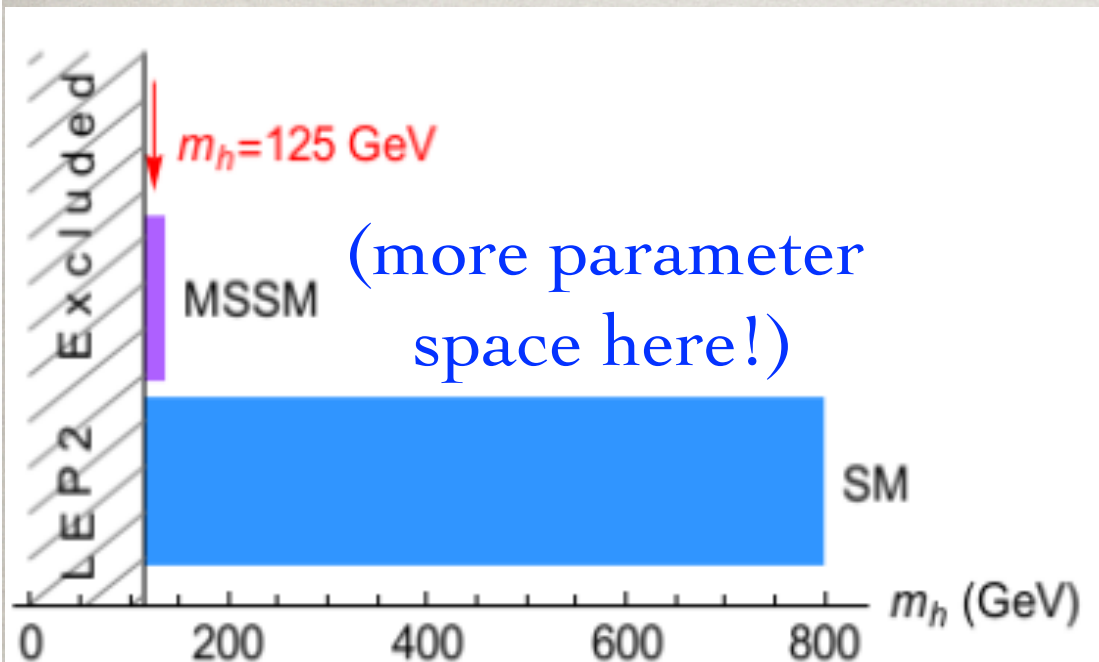
$$m_H^2 \approx M_Z^2 \cos^2 2\beta + \Delta m^2_{\text{SUSY}}$$

Tree-level $< (80 \text{ GeV})^2$ + loop-level: $> (45 \text{ GeV})^2$

→ Need large $\tan\beta$;

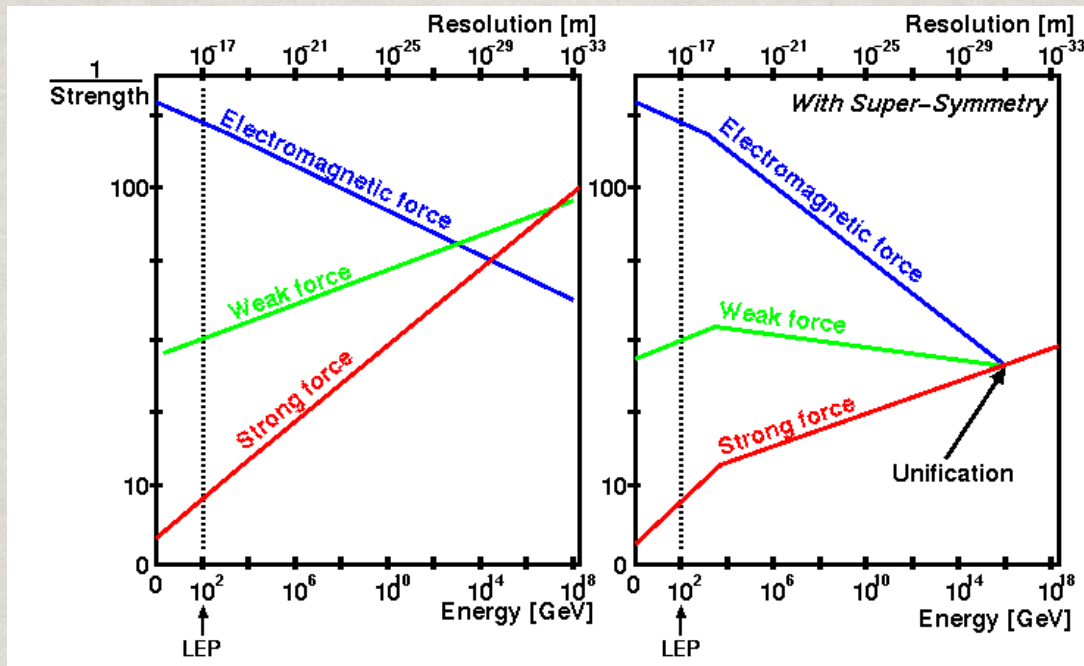
m_{stop} & mixing $X_t \gg m_t$

sensitive to μ , the Higgsino parameter

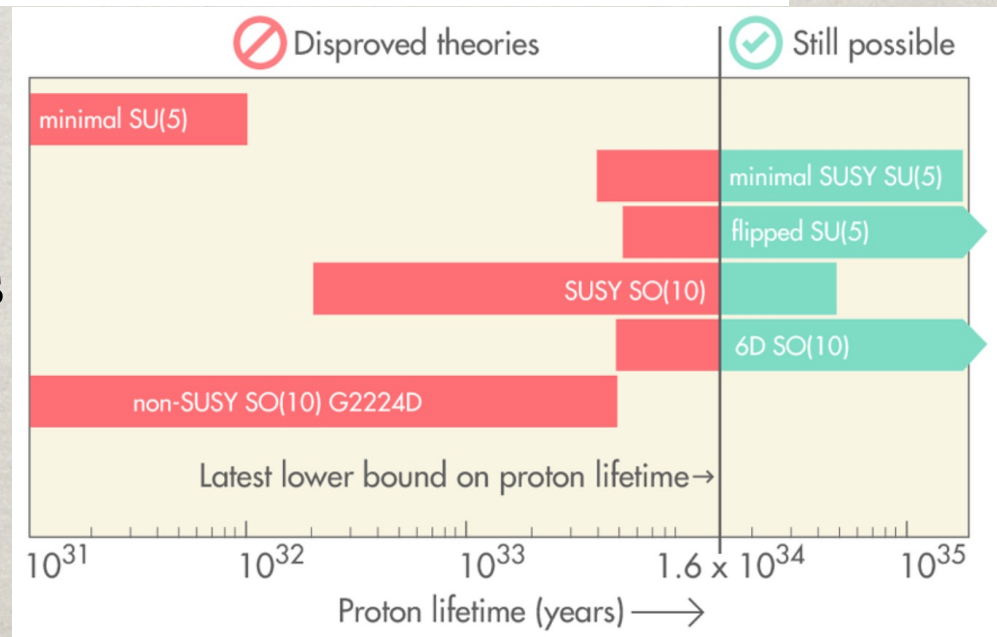


“Little hierarchy problem”

SUSY Grand Unification



Proton decay bounds
 --- SUSY helps!



SUSY breaking/mediation scenarios:

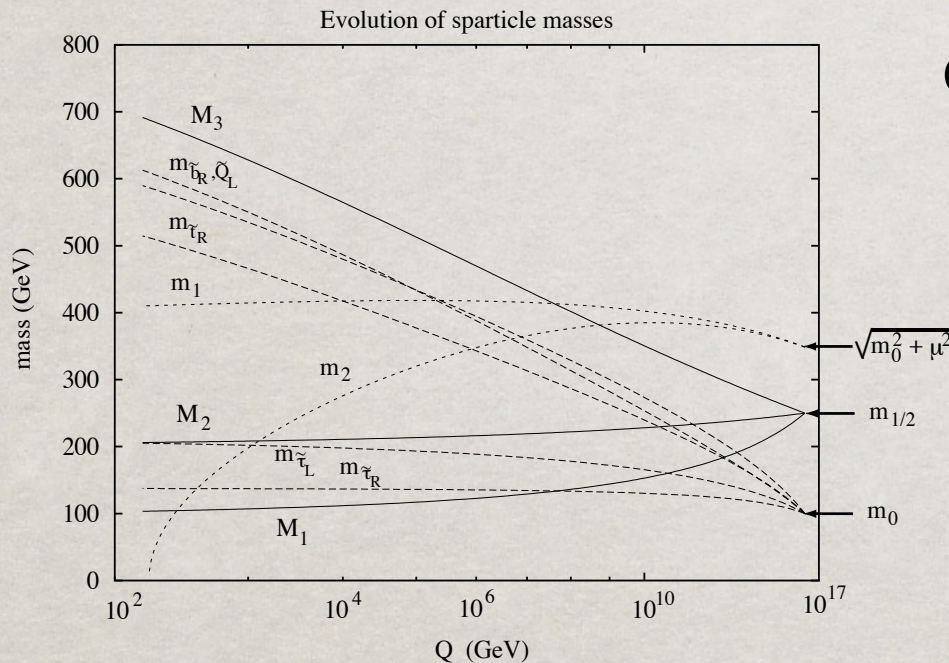
(*) "Minimal Super-gravity" (mSUGRA) scenario:

Gravity mediates the SUSY breaking from M_{GUT} to M_{EW} .

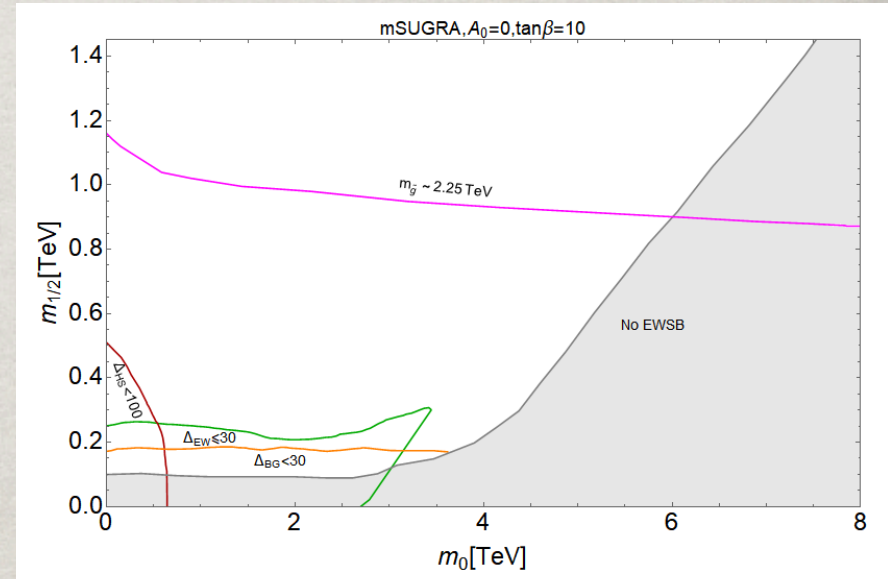
At the high scale: m_0 , $m_{1/2}$, A , $\tan\beta$, and $sign(\mu)$

However,

Currently, $\Delta_{BG}(mSUGRA) > 1000$



All masses at the EW scale predicted.



mSUGRA/CMSSM highly unnatural under all measures: ruled out by LHC!

(*) “Gauge mediation” scenario: *
 parameters: Λ , M_m , N , $\tan\beta$, $\text{sign}(\mu)$

N : no. of messengers,
 M_m (10 – 100 TeV) :
 messenger mass scale,
 $\Lambda = F/M_m$ (10 – 100 TeV):
 SUSY breaking scale.

Spin- $\frac{1}{2}$ Goldstino LSP:

$$m_{\tilde{G}} = \frac{F}{\sqrt{3}M_{pl}} \approx \left(\frac{\sqrt{F}}{100 \text{ TeV}} \right)^2 \text{ eV}$$

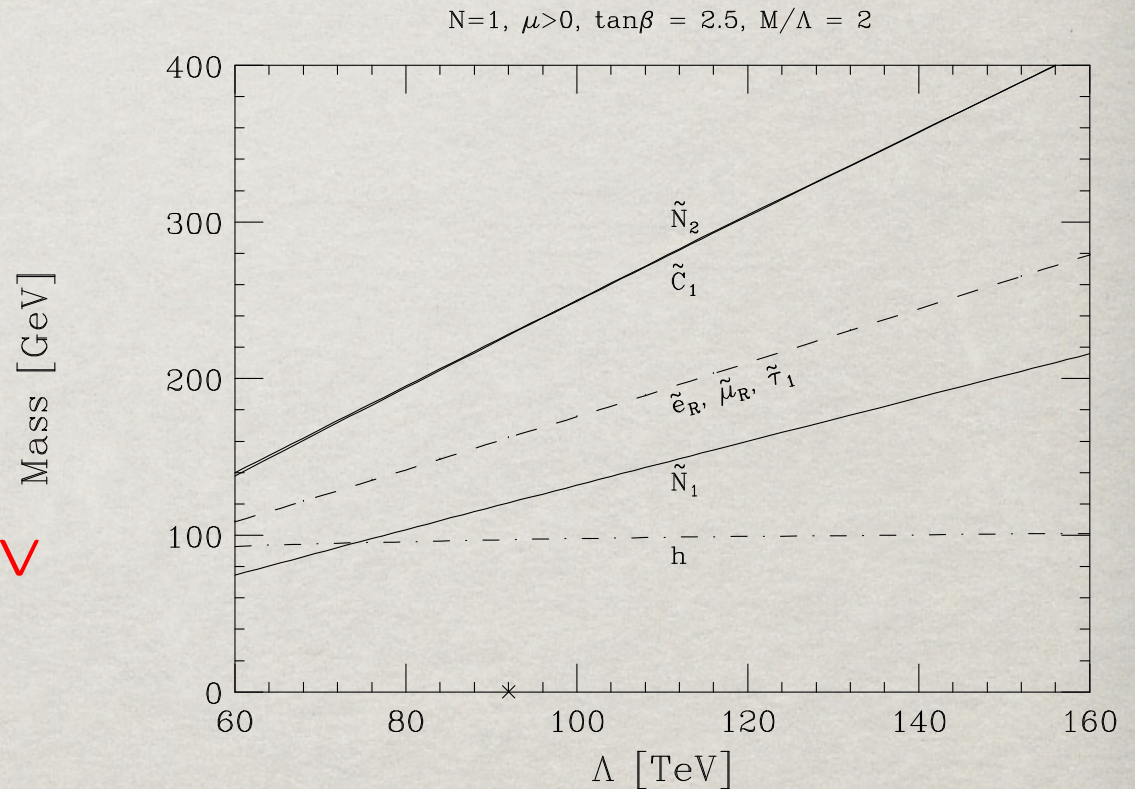
leading to missing energy.

Sparticle masses:

$$\tilde{M}_{\tilde{\chi}} \sim \kappa N \frac{\alpha_a}{4\pi} \Lambda,$$

$$\tilde{m}_{\tilde{f}}^2 \sim \kappa' N \sum_i C_i \frac{\alpha_i}{4\pi} \Lambda^2.$$

$$\tilde{e} \rightarrow e^+ \chi_1^0, \quad \chi_1^0 \rightarrow \gamma \tilde{G}$$



(*) “Anomaly mediation” scenario: †

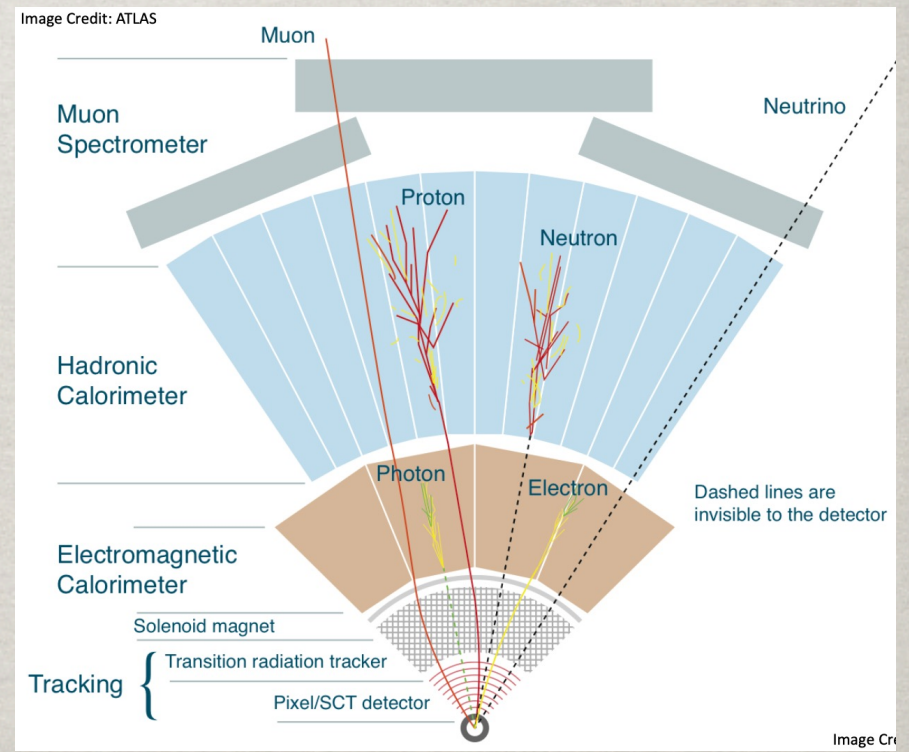
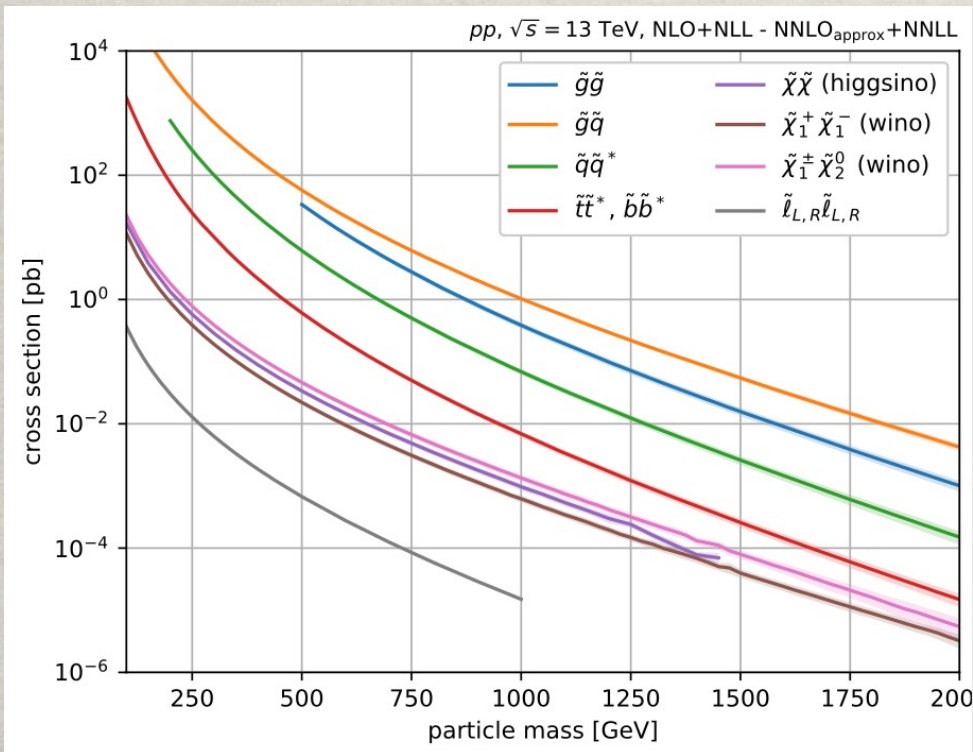
Scalar and gaugino masses generated by one-loop Super-Weyl anomaly (Gravity as bulk messengers).

(*) “Gaugino mediation” scenario: †

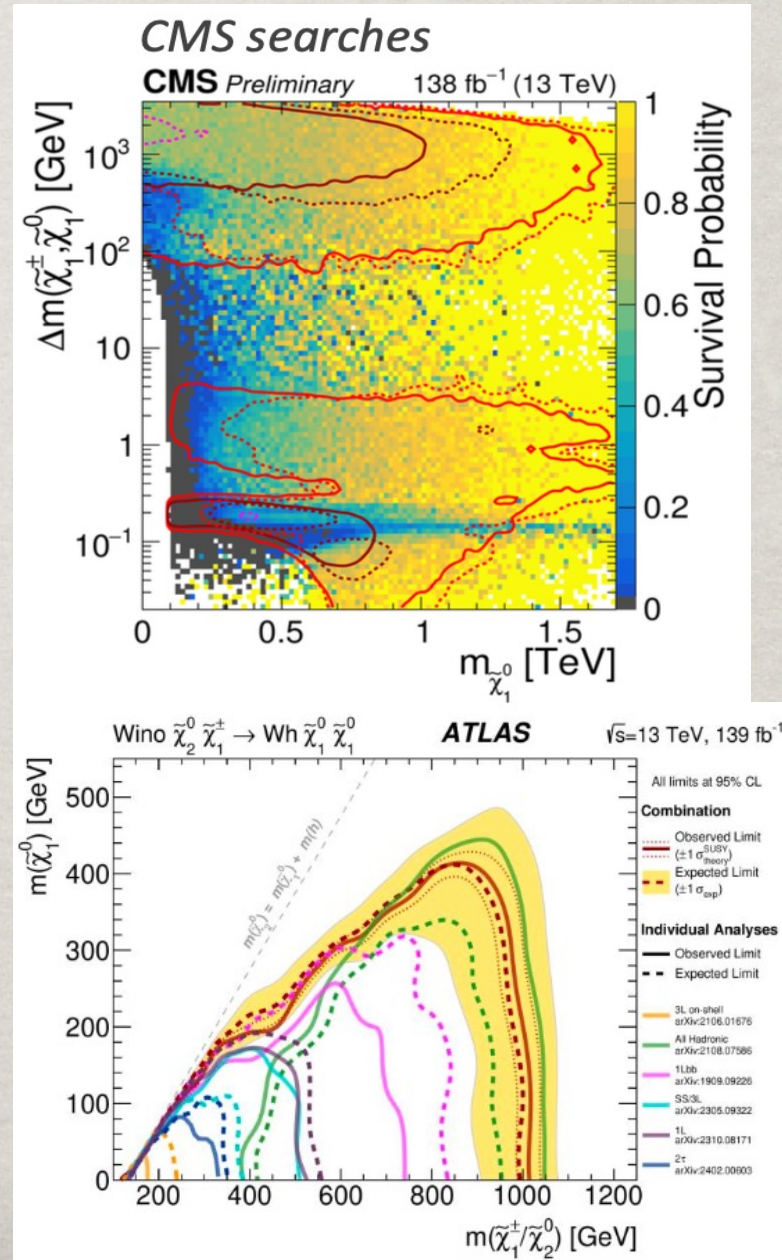
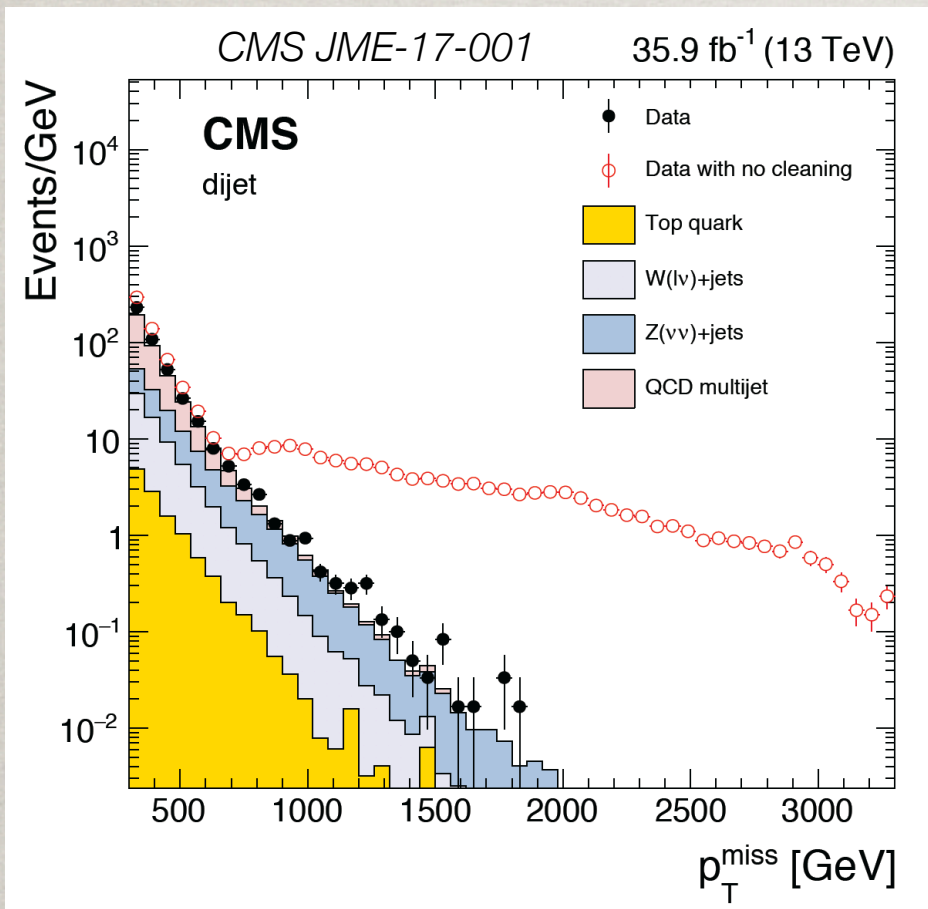
Gauginos as bulk messengers.

(*) “Mixed Modulus-Anomaly mediation” scenario: ‡

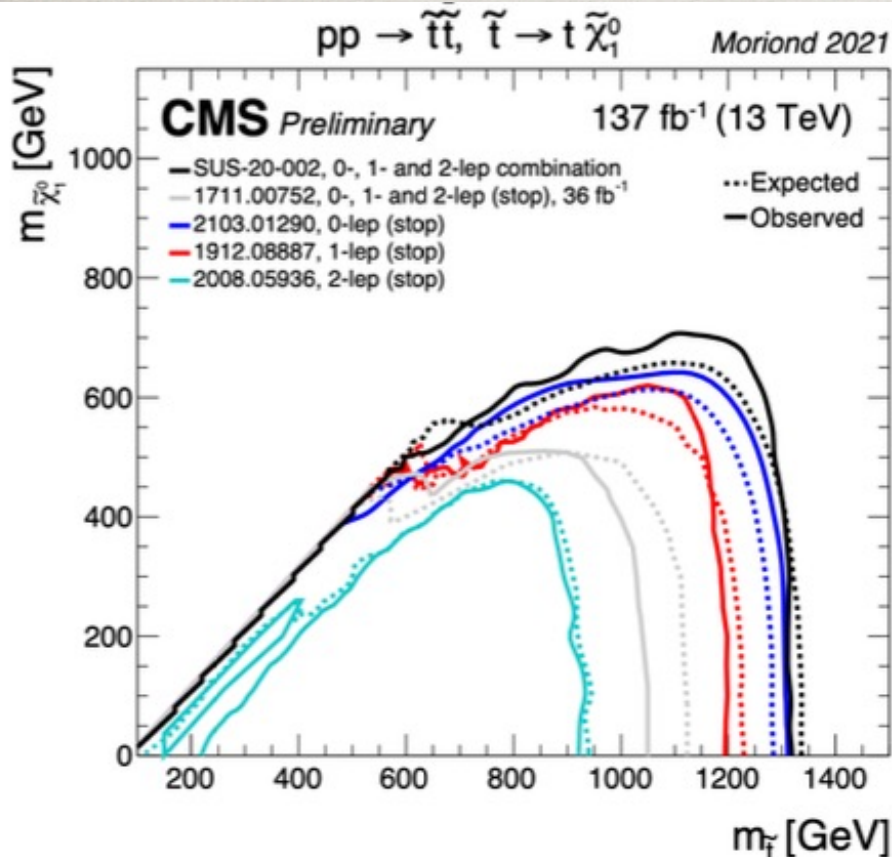
Only experiments can tell: A real challenge!



Smoking gun signal for SUSY --- Large missing transverse momentum:



Top-partner is the most likely suspect SUSY on Trial:



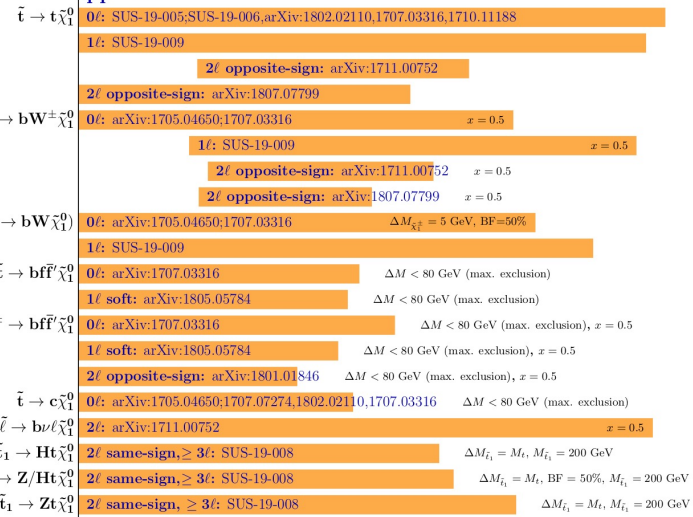
CMS (preliminary)

May 2019

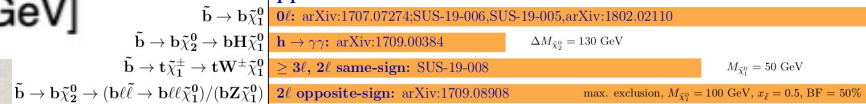
Overview of SUSY results: squark pair production

36/137 fb⁻¹ (13 TeV)

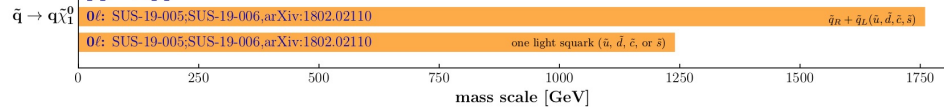
pp → $\tilde{t}\tilde{t}^*$



pp → $\tilde{b}\tilde{b}^*$

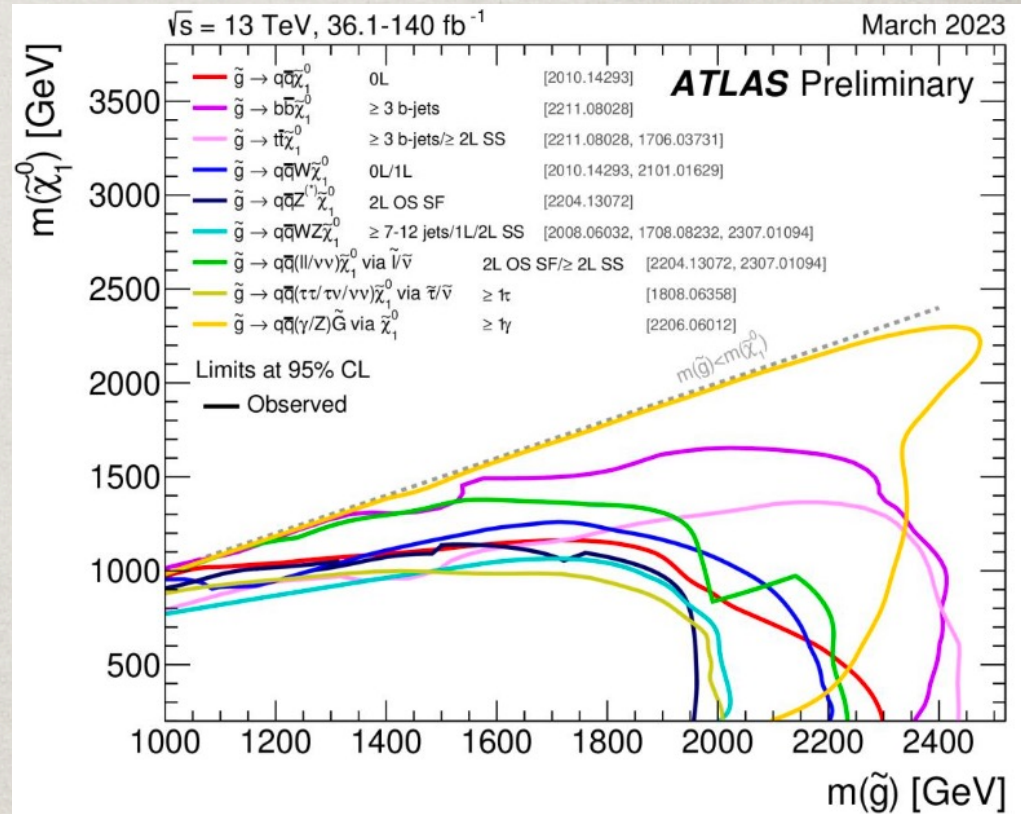
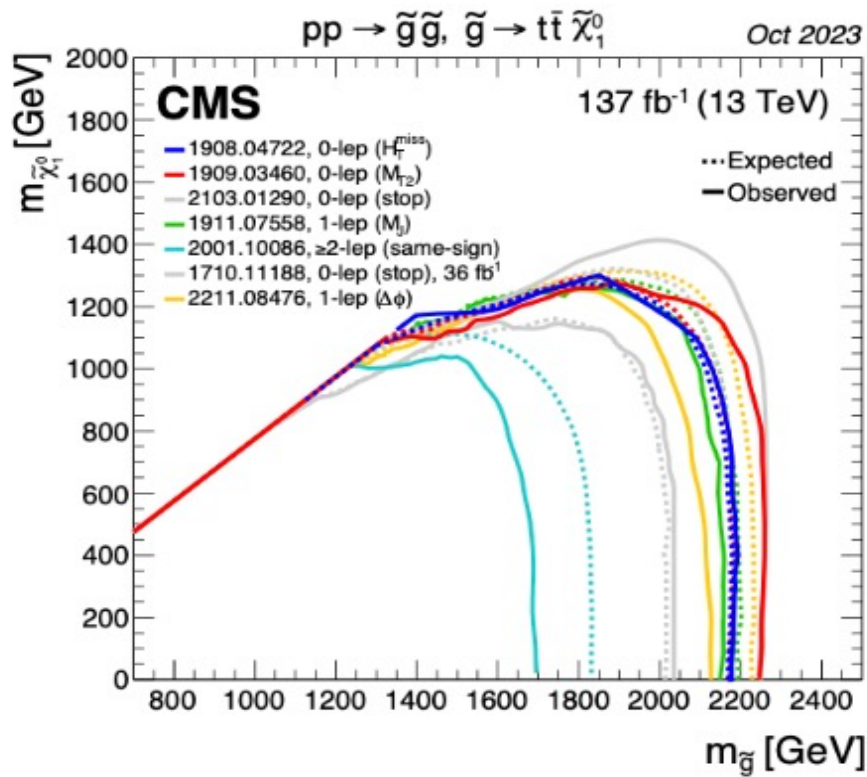


pp → $\tilde{q}\tilde{q}^*$



Selection of observed limits at 95% C.L. (theory uncertainties are not included). Probe up to the quoted mass limit for light LSPs unless stated otherwise. The quantities ΔM and x represent the absolute mass difference between the primary sparticle and the LSP, and the difference between the intermediate sparticle and the LSP relative to ΔM , respectively, unless indicated otherwise.

Gluinos are the next candidate!



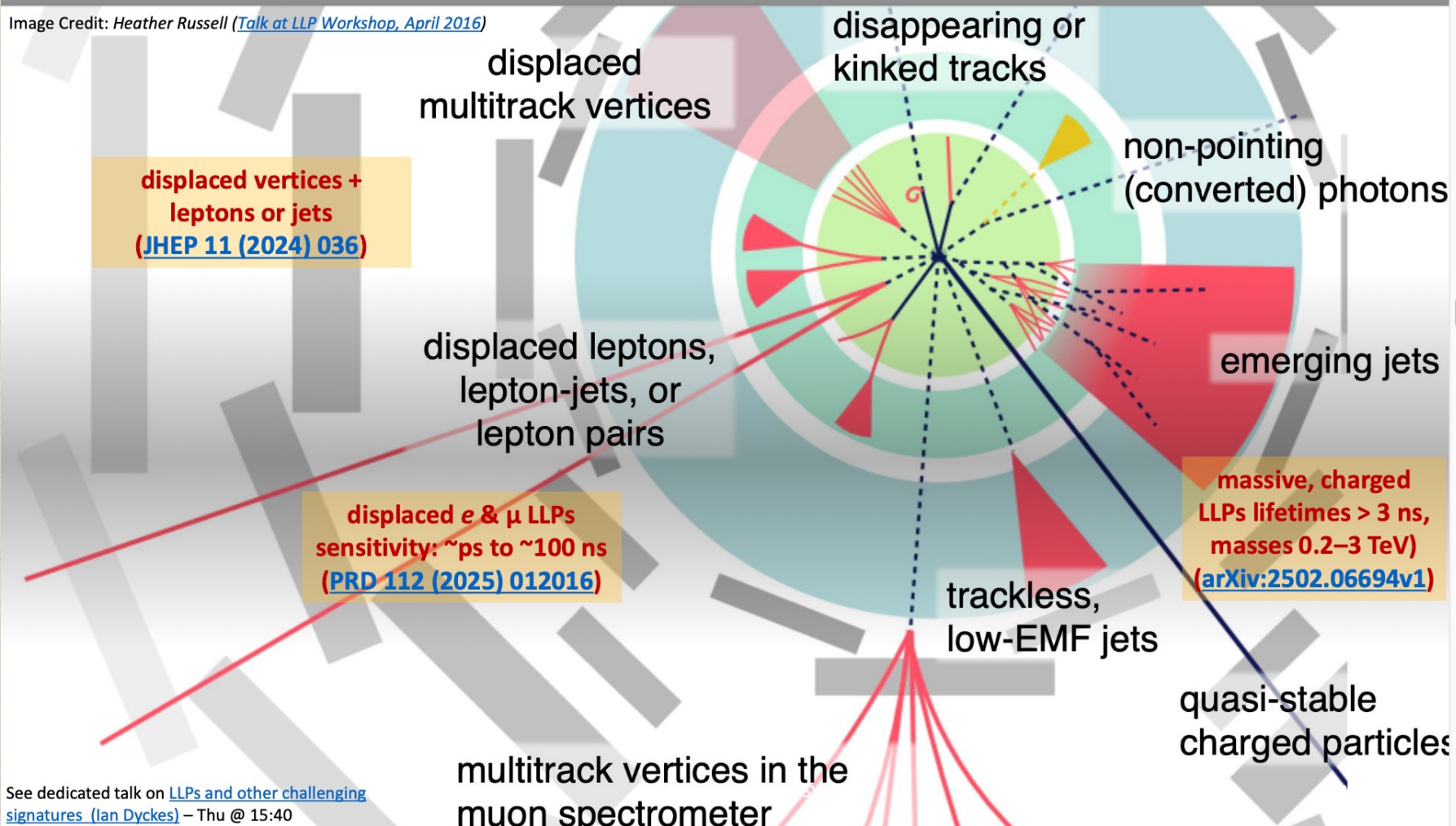
Long-lived particles (LLP)

18 August 2025

ATLAS Searches for SUSY

22

Image Credit: Heather Russell ([Talk at LLP Workshop, April 2016](#))



See dedicated talk on [LLPs and other challenging signatures](#) (Ian Dyckes) – Thu @ 15:40

Missing siblings:

Two Higgs-Doublet Model

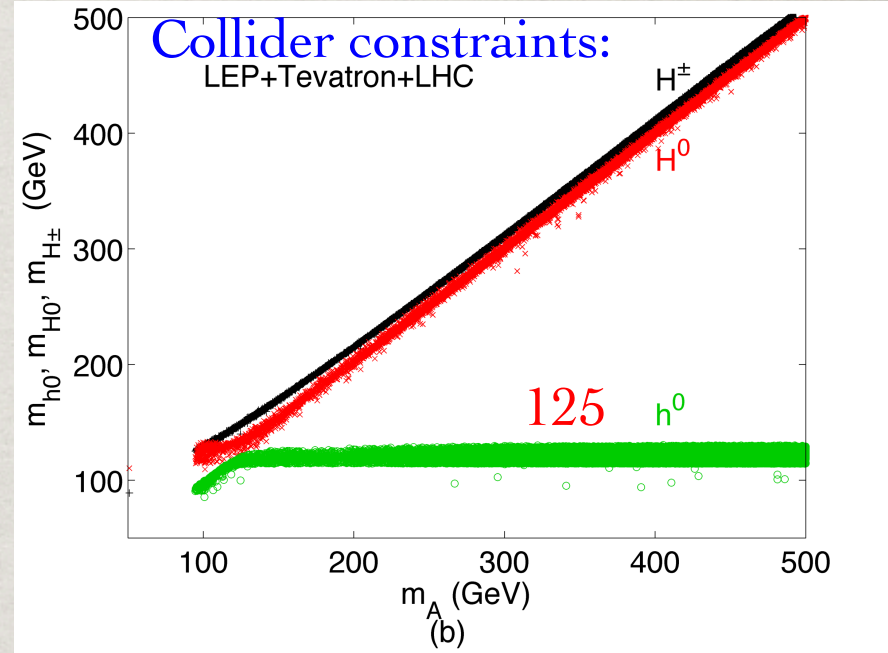
5 Higgs bosons: h^0, H^0, A^0, H^\pm

Tree-level mass given by:

$$M_A, \tan\beta$$

General 2HDM with Z_2 symm

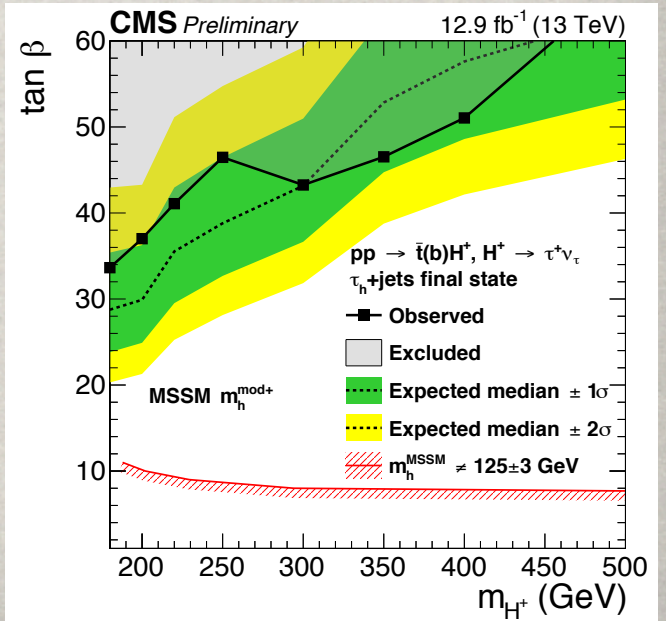
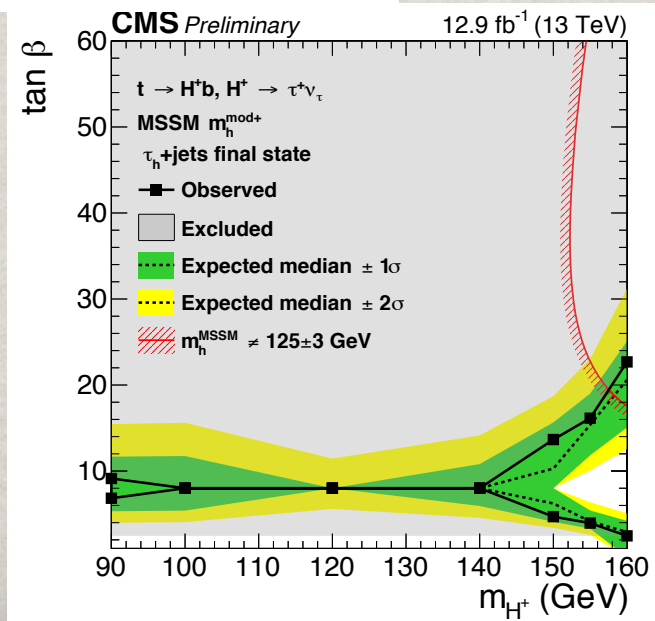
Types	Φ_1	Φ_2	u_R	d_R	ℓ_R	Q_L, L_L	Φ_1	Φ_2
Type-I	+	-	-	-	-	+		u, d, ℓ
Type-II	+	-	-	+	+	+	d, ℓ	u
Type-L	+	-	-	-	+	+	ℓ	$u, d,$
Type-F	+	-	-	+	-	+	d	u, ℓ



$$\begin{pmatrix} H^\pm \\ G^\pm \end{pmatrix} = \begin{pmatrix} c_\beta & -s_\beta \\ s_\beta & c_\beta \end{pmatrix} \begin{pmatrix} \phi_2^\pm \\ \phi_1^\pm \end{pmatrix}$$

$$\begin{pmatrix} A^0 \\ G^0 \end{pmatrix} = \begin{pmatrix} c_\beta & -s_\beta \\ s_\beta & c_\beta \end{pmatrix} \begin{pmatrix} G_2^0 \\ G_1^0 \end{pmatrix}$$

$$\begin{pmatrix} h \\ H \end{pmatrix} = \begin{pmatrix} c_\alpha & -s_\alpha \\ s_\alpha & c_\alpha \end{pmatrix} \begin{pmatrix} \phi_2^0 \\ \phi_1^0 \end{pmatrix}$$



“Naturally speaking”:

H^0 should not be a lonely particle; has an
“interactive friend circle”: t, W^\pm, Z
and partners $\tilde{t}, \tilde{W}^\pm, \tilde{Z}, \tilde{H}^{\pm,0} \dots$

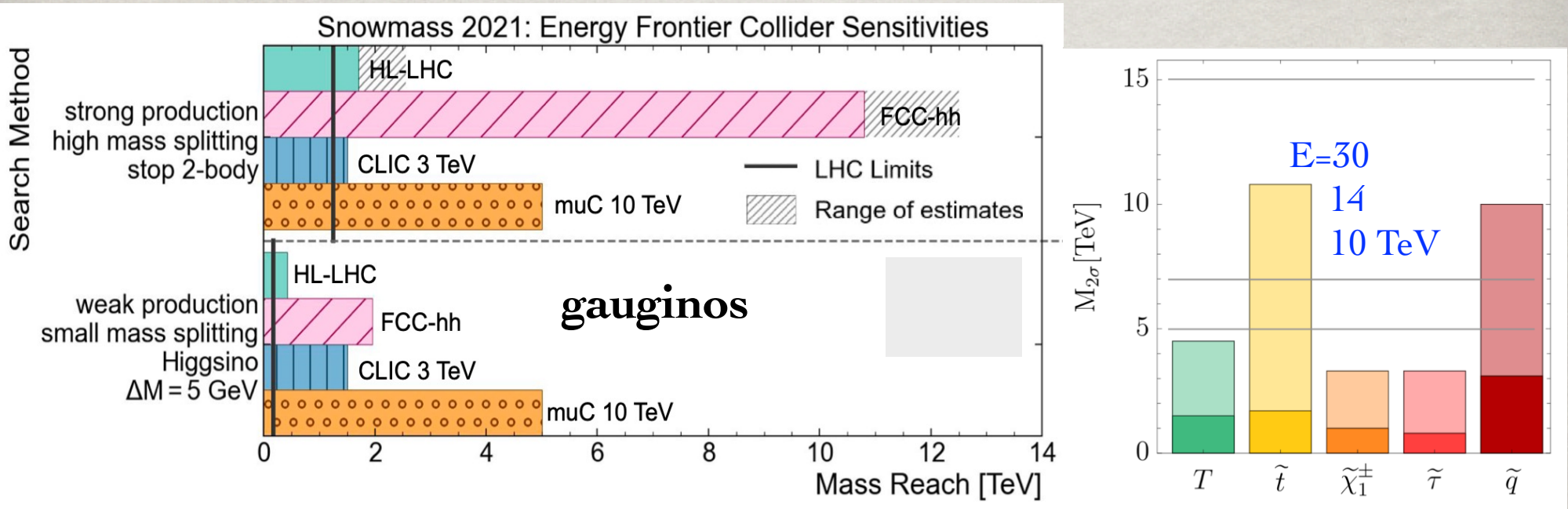
“Little hierarchy problem”



The Jury is still out (again) ...

Future prospects: pushing the “Naturalness” limit

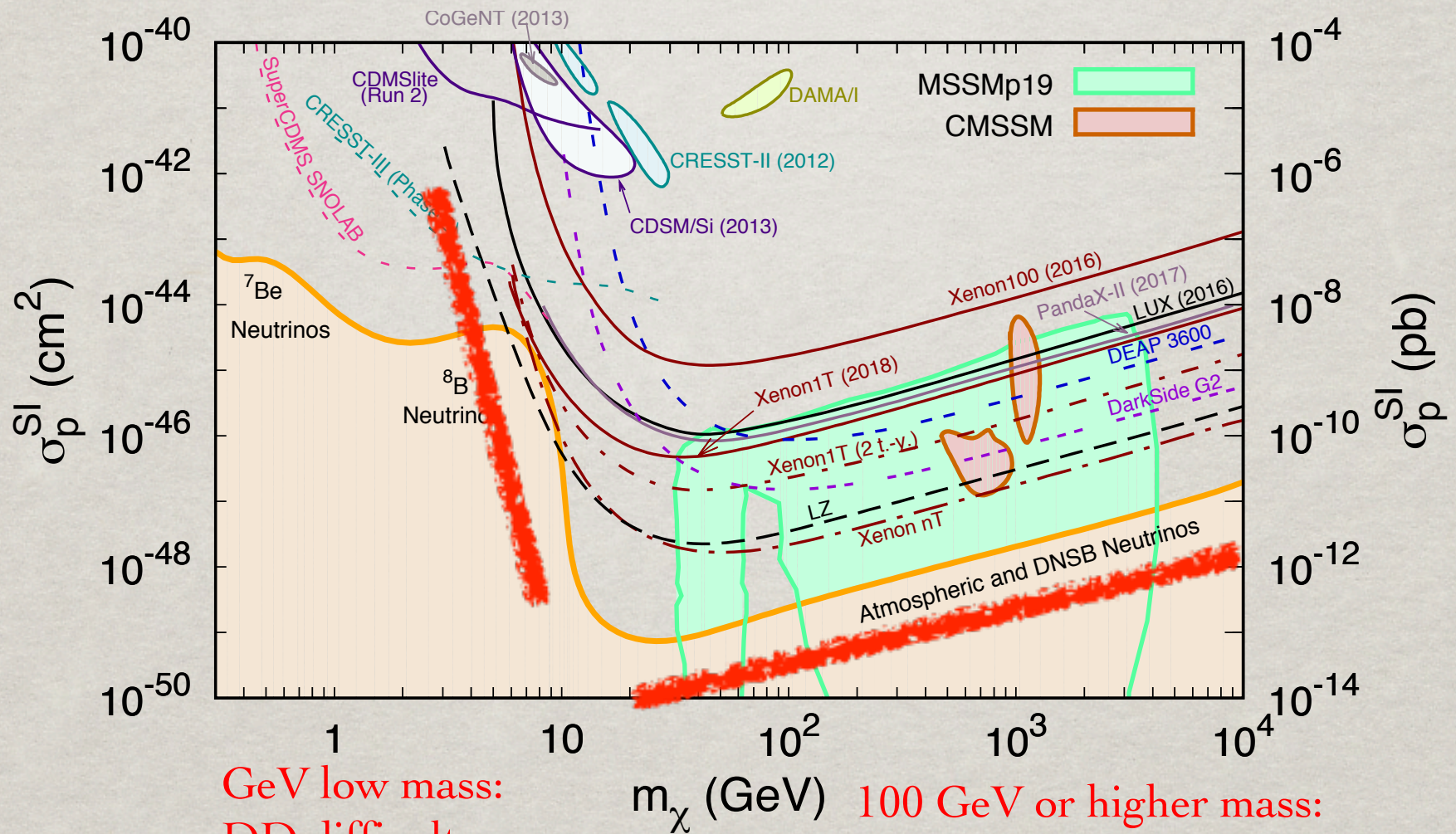
The searches for top quark partners & gluinos, gauginos ...



→ Higgs mass fine-tune: $\delta m_H/m_H \sim 1\% (1 \text{ TeV}/\Lambda)^2$
 Thus, $m_{\text{stop}} > 8 \text{ TeV} \rightarrow 10^{-4}$ fine-tune!

SUSY WIMP DM

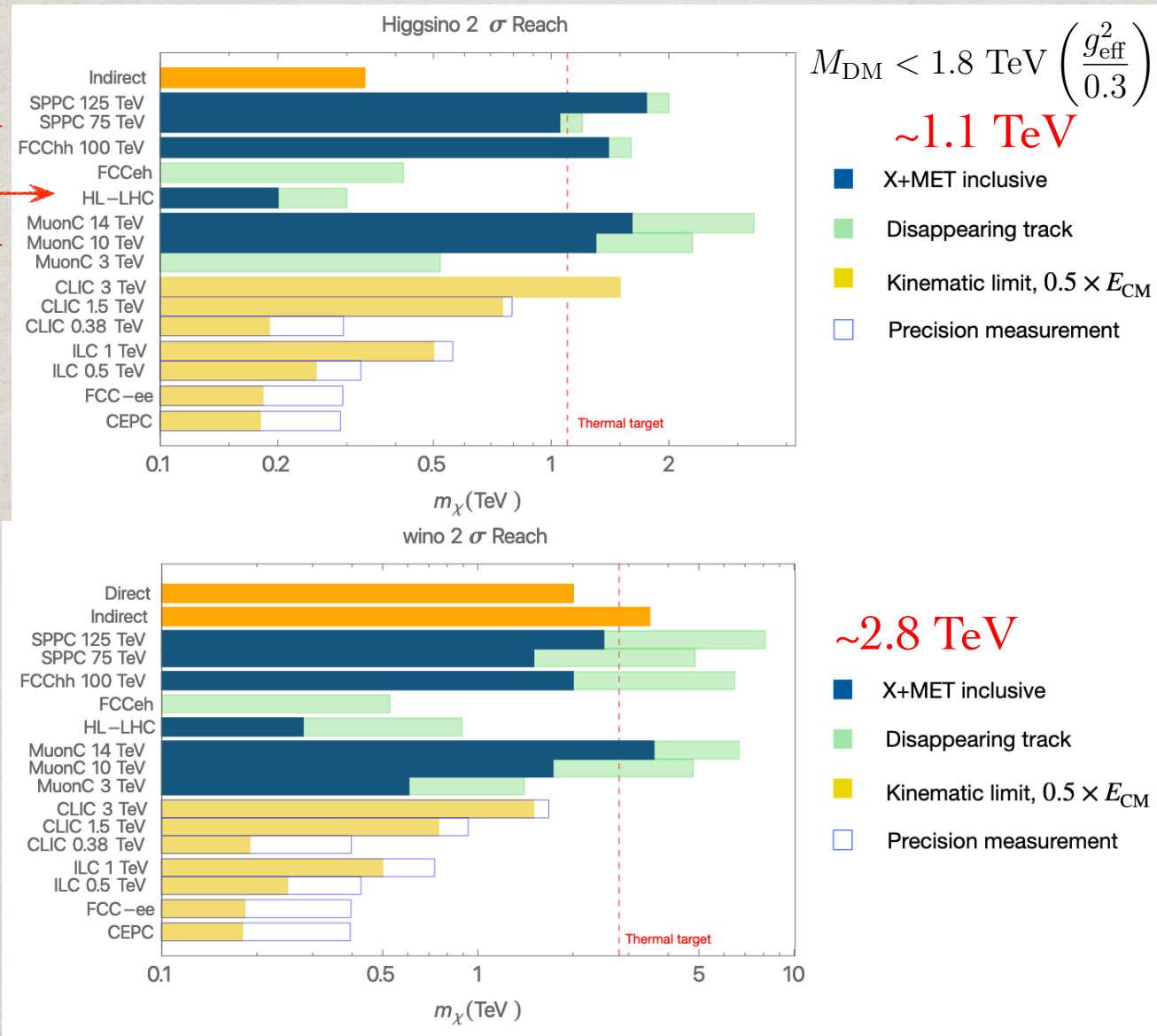
The lightest neutralino could be a natural WIMP dark matter candidate!



GeV low mass:
DD difficult;
Collider complementary

100 GeV or higher mass:
DD + ID + HE Collider

Utilizing the missing-mass technique for DM search, covering the thermal target

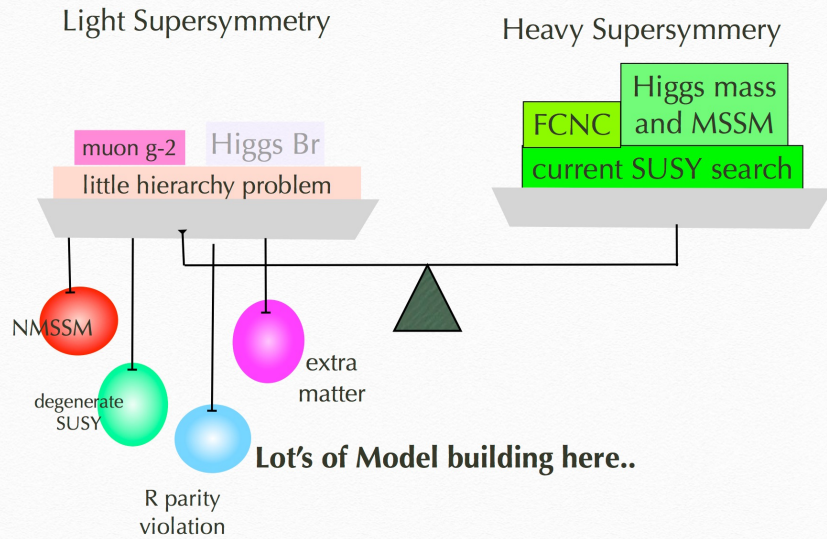


TH, Z. Liu, L.T. Wang, X. Wang: arXiv:2009.11287; arXiv:2203.07351

Alternative to “Naturalness”: SUSY “splitted”?

Nima Arkani-Hamed,
Giudice, Delgado, ...

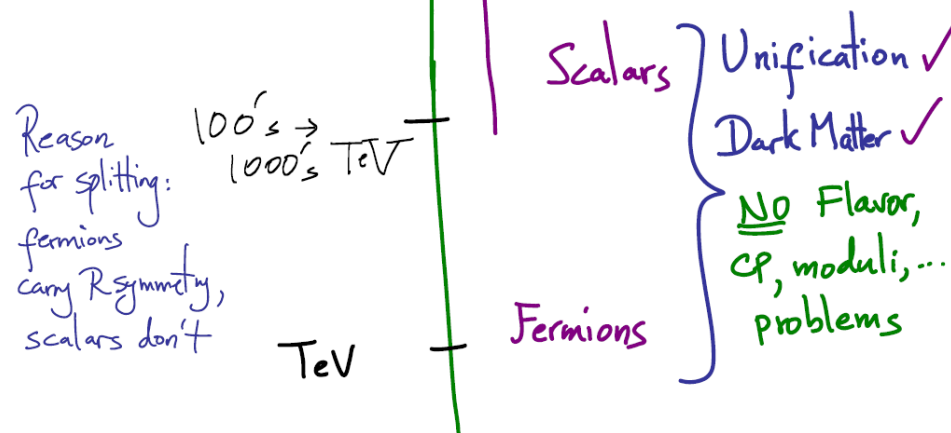
Mind of SUSY theorists



Split Supersymmetry: Philosophy

- SUSY is irrelevant to the hierarchy problem
 - Cosmological constant problem suggests fine-tuning mechanism → may also apply to the gauge hierarchy
- Break SUSY at the GUT scale
 - Scalars become ultra-heavy (except 1 light Higgs): $m_s \sim 10^{9-12}$ GeV
 - Fermions protected by chiral symmetry
- Phenomenological Successes:
 - Retain gauge coupling unification
 - Higgs mass predicted to be 'heavier': $m_H \sim 120-150$ GeV
 - Flavor & CP problems are automatically solved
 - Proton decay is delayed (occurs via dimension-6 operator)
- Collider signatures & Dark Matter implications substantially different!

SPLIT SUSY



Alternatives to “Naturalness”?

- Anthropic Landscape– not a *dynamical* explanation

String theory predicts: there are at least $10^{272,000}$ vacua!

And we just so happen to live here ... (Don't ask why!)

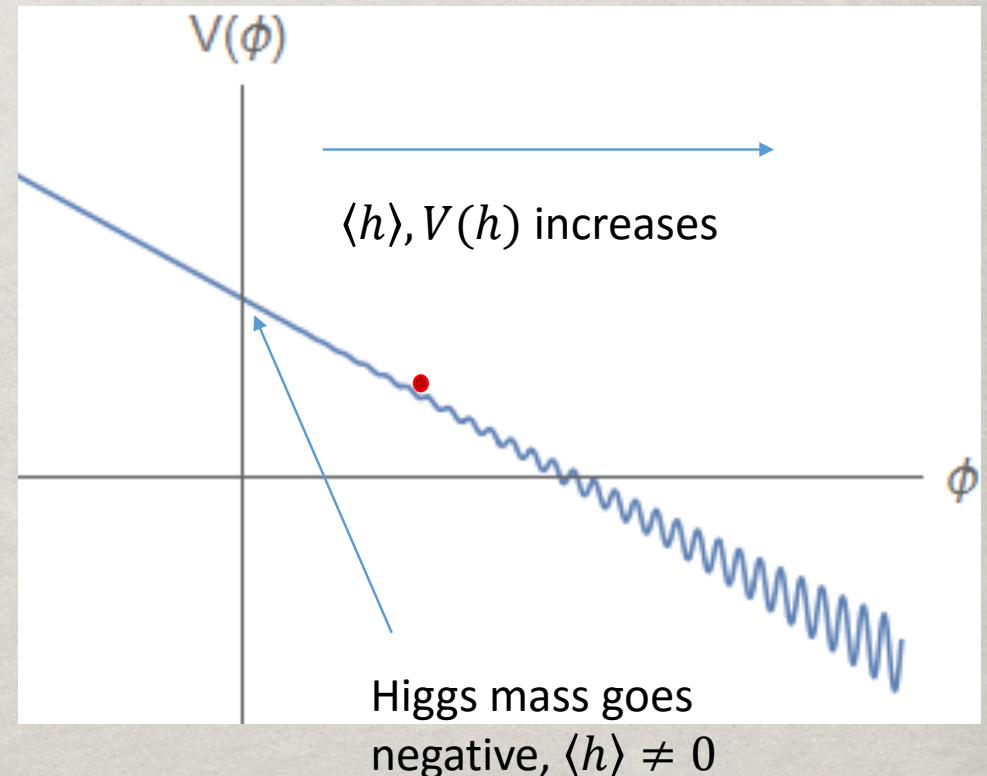
- Cosmological Relaxation (Graham, Kaplan, Rajendran 2015)

$$\mathcal{L} \supset g\Lambda^3\phi - (M^2 - g\Lambda\phi) |h|^2$$

Λ is the EFT cutoff scale; “bare” Higgs mass is $M \sim \Lambda$

Technically natural-- shift symmetry for ϕ broken only by small coupling g

ϕ must vary over a range $\Delta\phi \sim \Lambda/g$ to scan the Higgs mass by $O(1)$



“...settling the ultimate fate of naturalness is perhaps the most profound theoretical question of our time”

-- Nima Arkani-Hamed

The central questions today are not details — but structural: origin of spacetime, UV/IR connection, standard model \rightarrow real theory

