

Super-Heavy Dark Matter

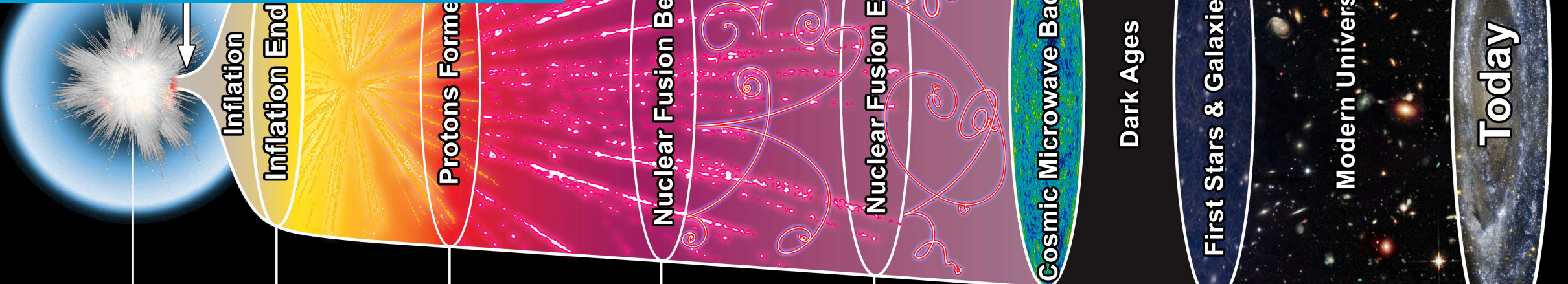
Tom Tong

Center for Particle Physics Siegen, University of Siegen

CPPS Retreat

15 February 2024

Radius of the Visible



What Could Dark Matter Be?

Mass, in electron volts (eV)



ULTRALIGHT DARK MATTER

Mass range
 $\sim 10^{-22}$ eV to $\sim 10^{-6}$ eV
Experiments
CASPEr, MAGIS-100



AXIONS

Mass range
 $\sim 10^{-6}$ eV to $\sim 10^{-3}$ eV
Experiments
ADMX, MADMAX,
QUAX, CAPP-8TB



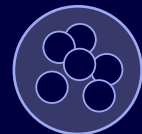
PRIMORDIAL BLACK HOLES

Mass range
 ~ 1 to ~ 30
solar masses
Experiments
LIGO/Virgo



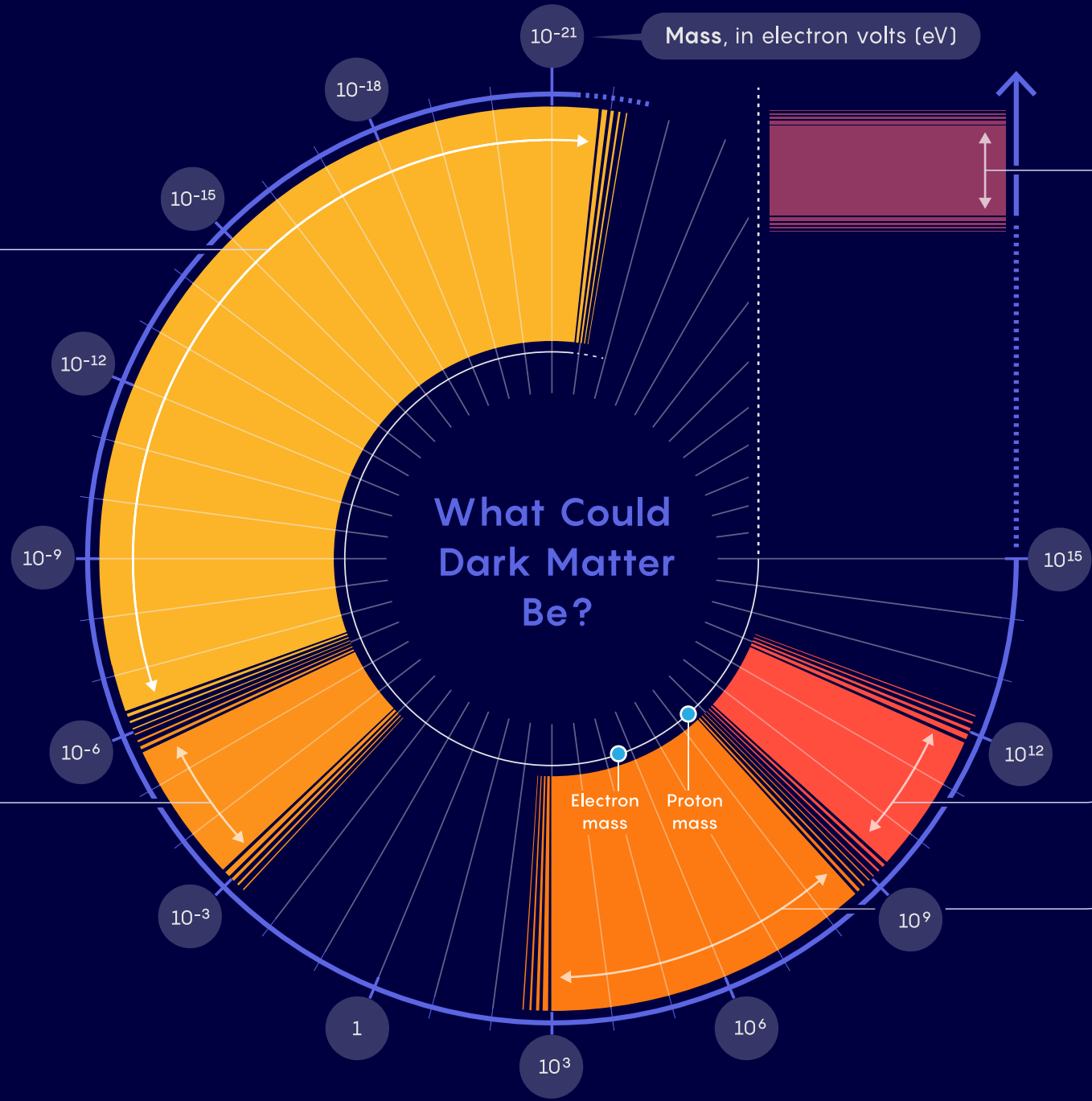
WIMPs

Mass range
 ~ 1 GeV to ~ 1 TeV
Experiments
XENONnT,
PandaX-4T,
LZ, CRESST, DAMA,
COSINE-100

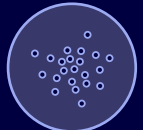


SUB-GeV DARK MATTER

Mass range
 ~ 1 keV to ~ 1 GeV
Experiments
SENSEI, TESSERACT

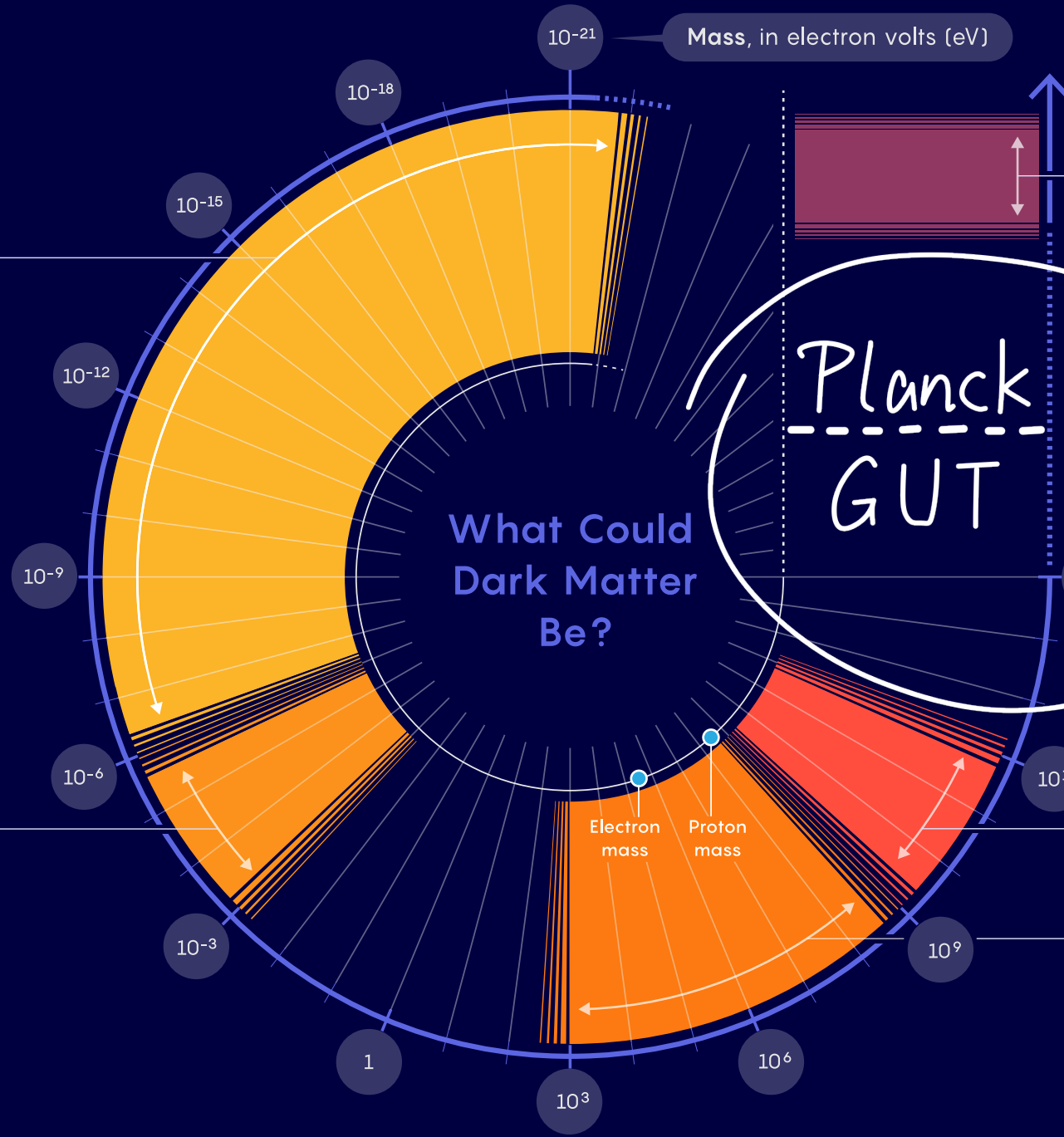


Mass, in electron volts (eV)



ULTRALIGHT DARK MATTER

Mass range
~ 10^{-22} eV to ~ 10^{-6} eV
Experiments
CASPEr, MAGIS-100



What Could Dark Matter Be?

Planck
GUT

10^{28}
 10^{25}

Electron mass
Proton mass

PRIMORDIAL BLACK HOLES

Mass range
~1 to ~30 solar masses
Experiments
LIGO/Virgo



WIMPs

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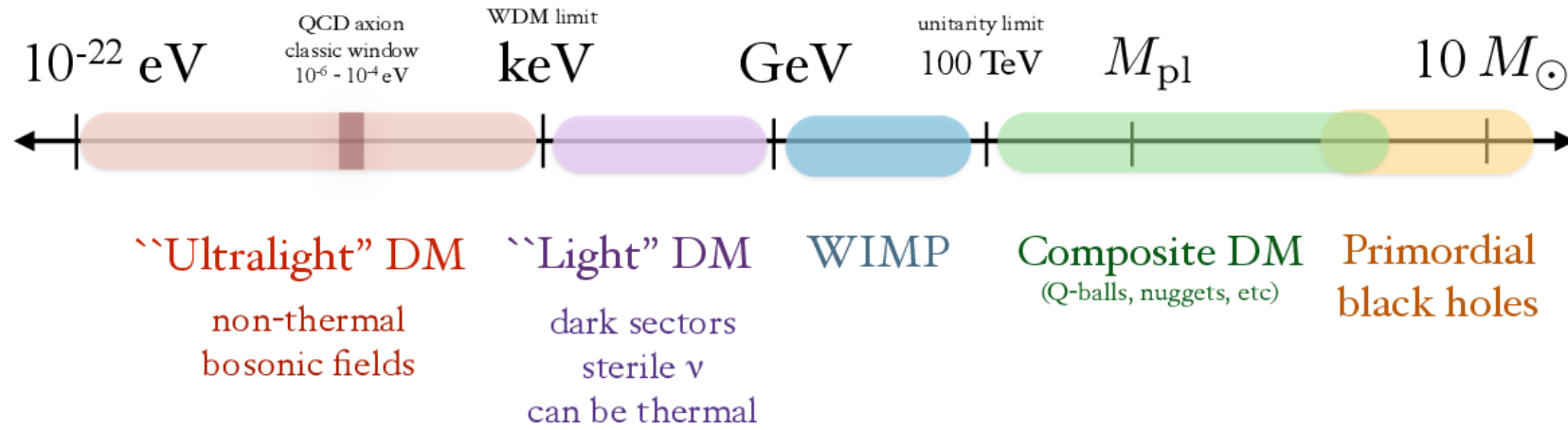
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ADMX, MADMAX, QUAX, CAPP-8TB



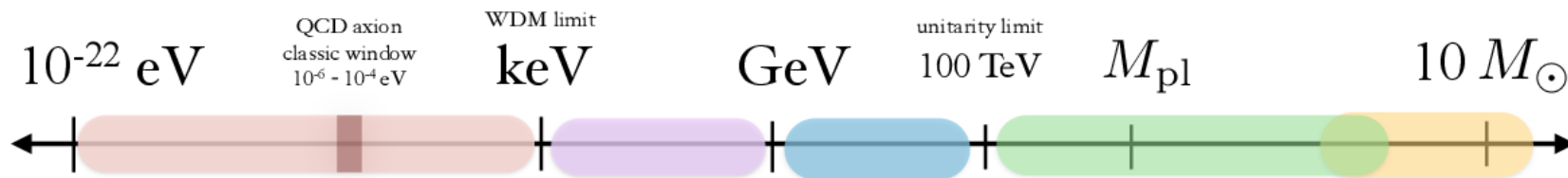
Mass scale of dark matter

(not to scale)



Mass scale of dark matter

(not to scale)



“Ultralight” DM

non-thermal
bosonic fields

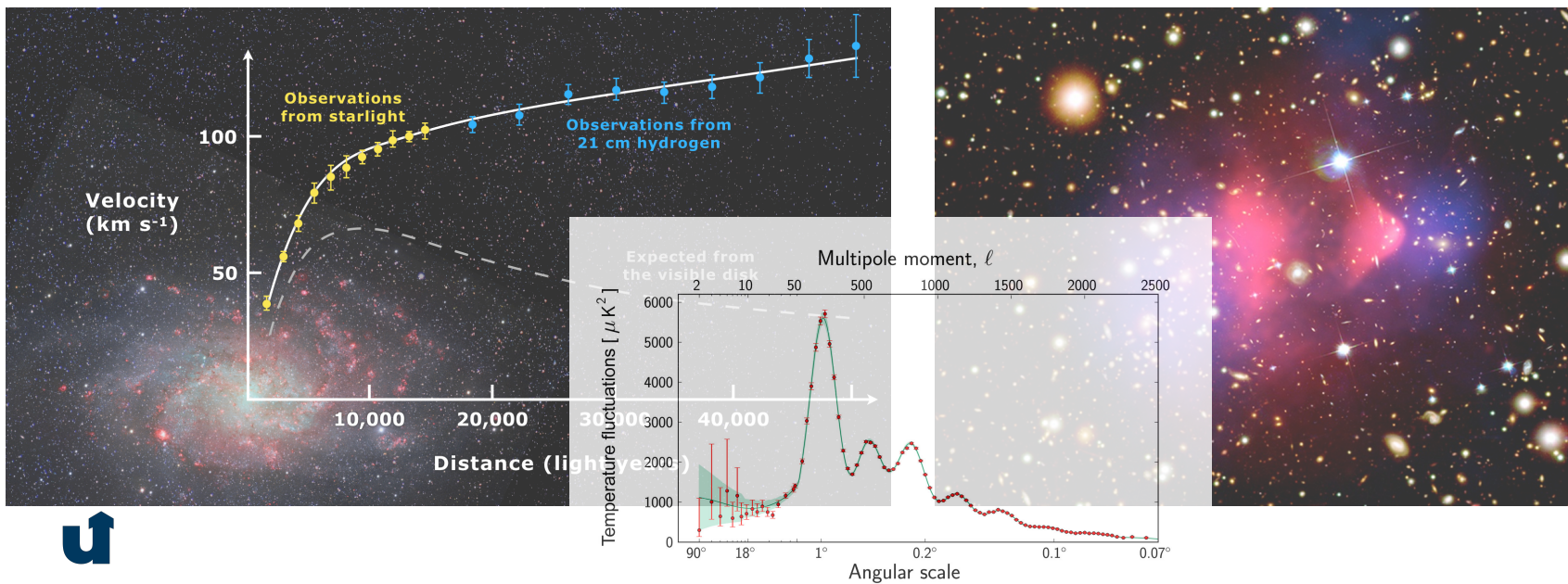
“Light” DM

dark sectors
sterile ν
can be thermal

WIMP

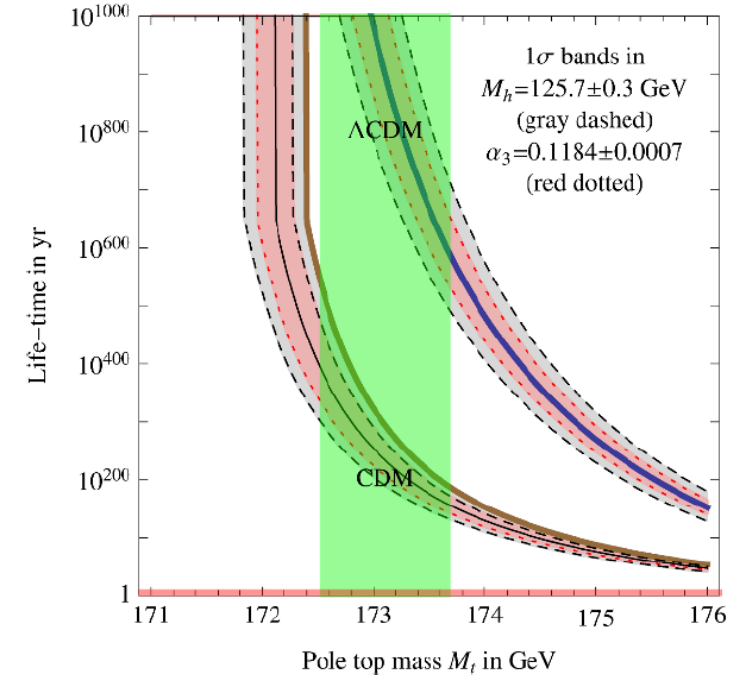
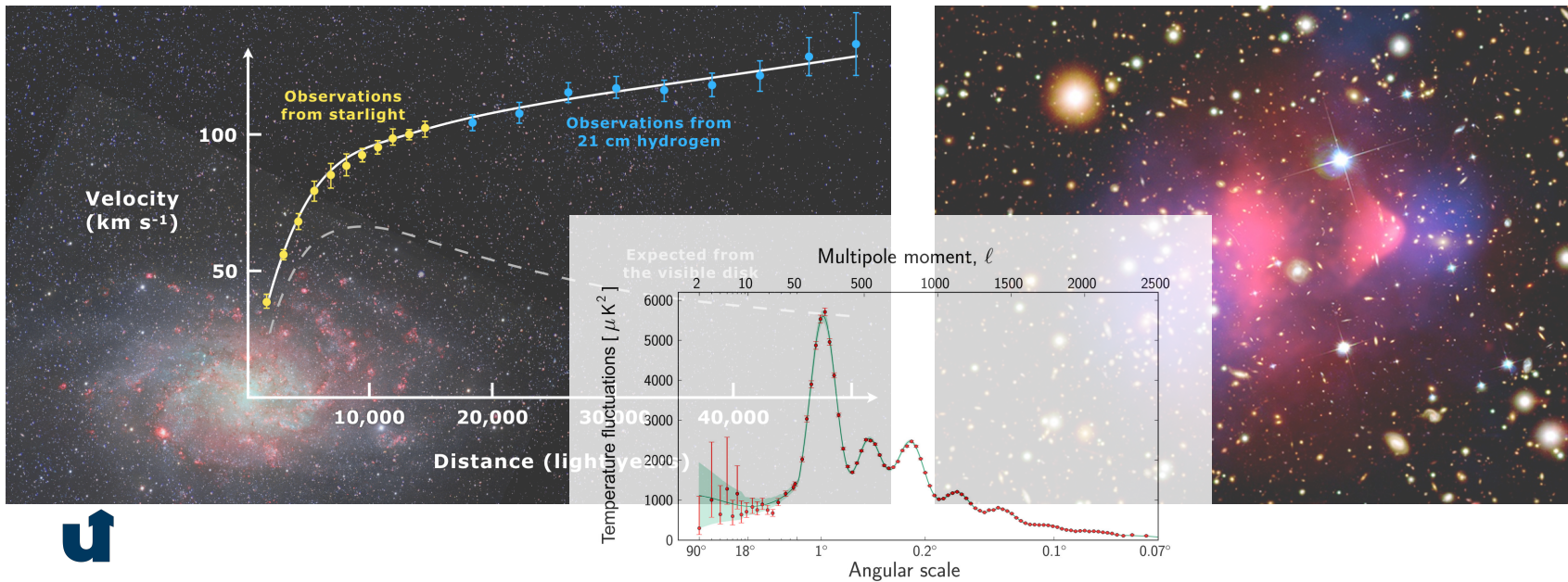
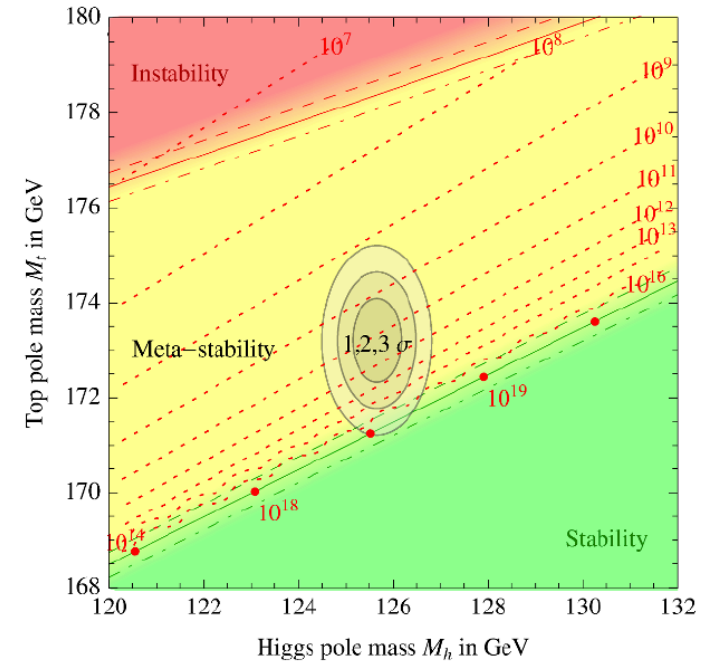
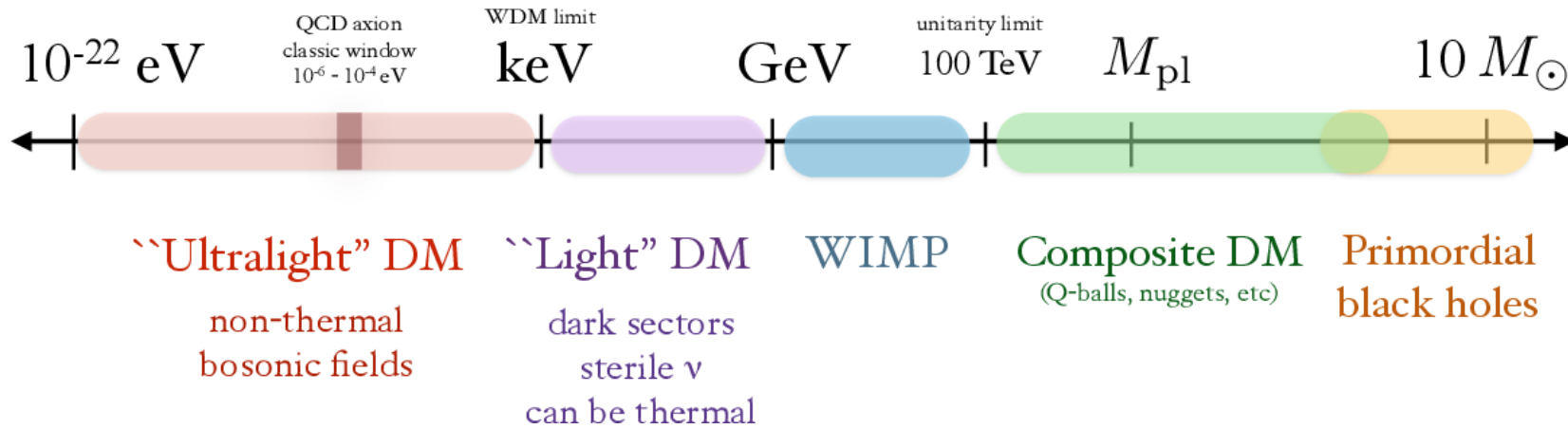
Composite DM
(Q-balls, nuggets, etc)

Primordial
black holes



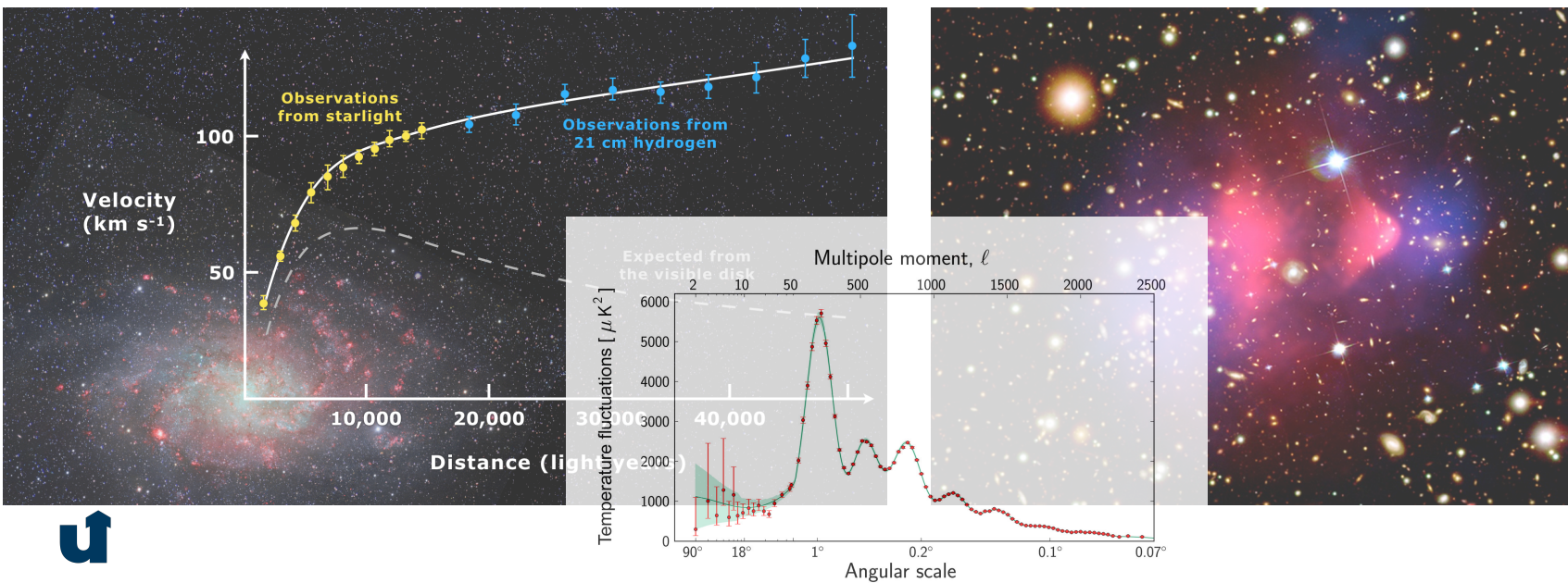
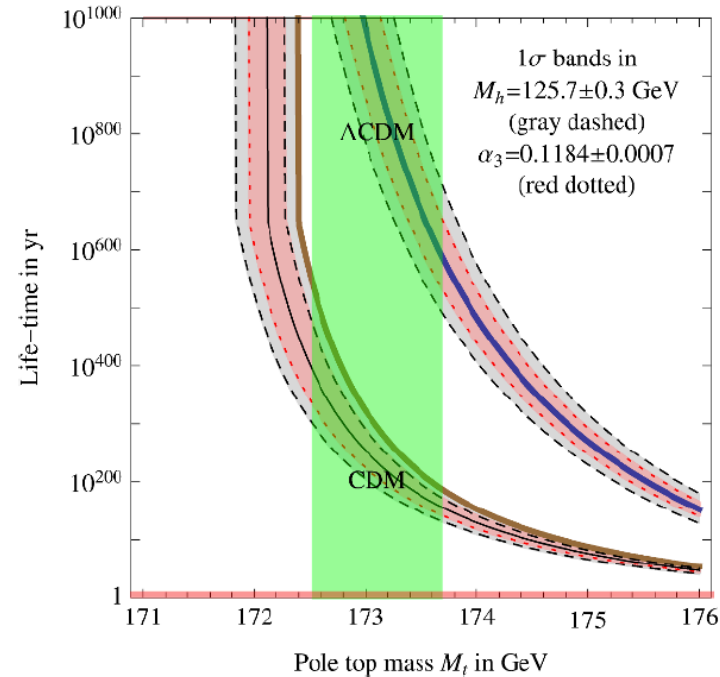
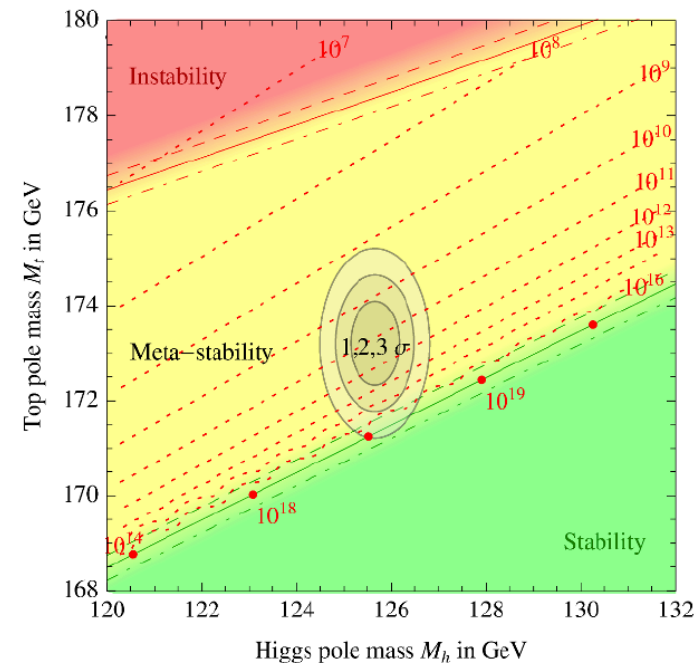
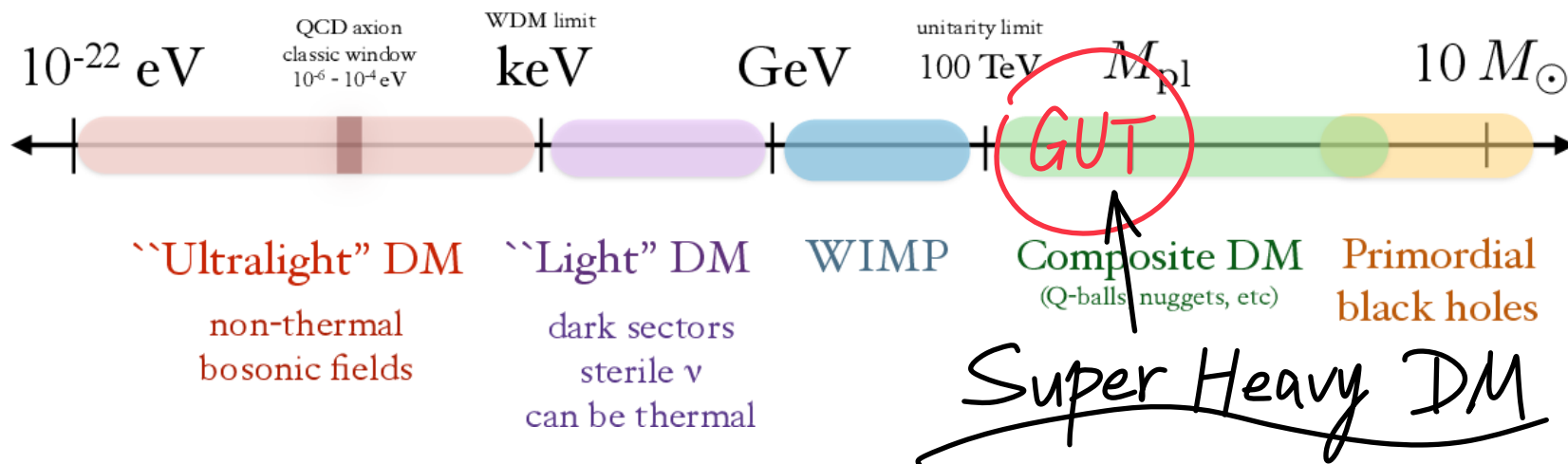
Mass scale of dark matter

(not to scale)



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Quiz: If SHDM is near the GUT scale, how do we probe it?

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Galaxy-Size Collider!

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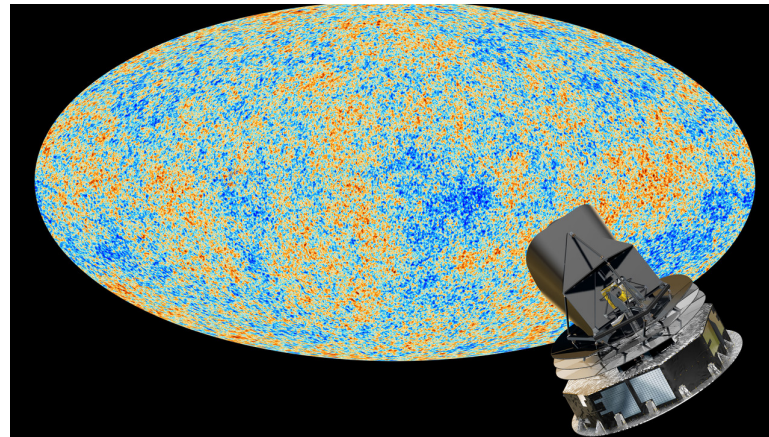
Galaxy-Size Collider!



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Galaxy-Size Collider!

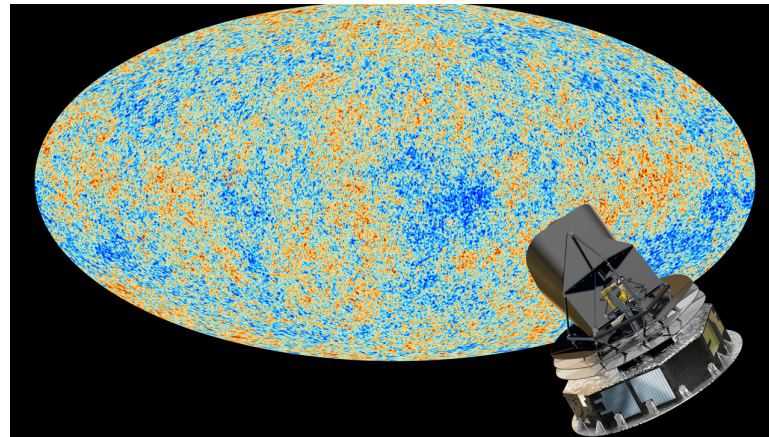


CMB Polarisation

Quiz: If SHDM is near the GUT scale, how do we probe it?



Galaxy-Size Collider!



CMB Polarisation

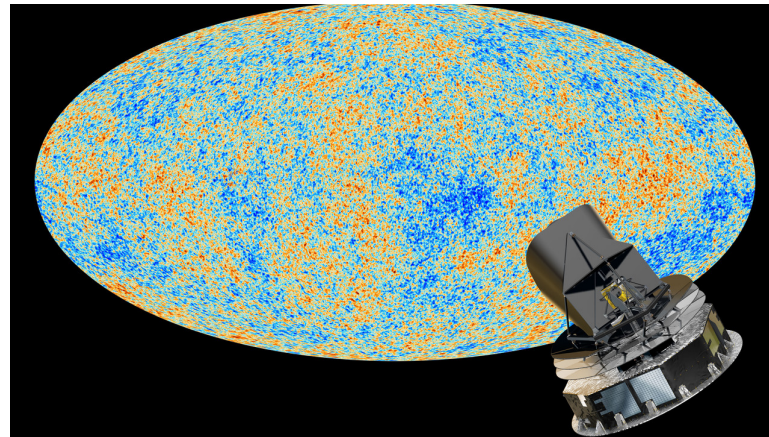


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Galaxy-Size Collider!

u



CMB Polarisation



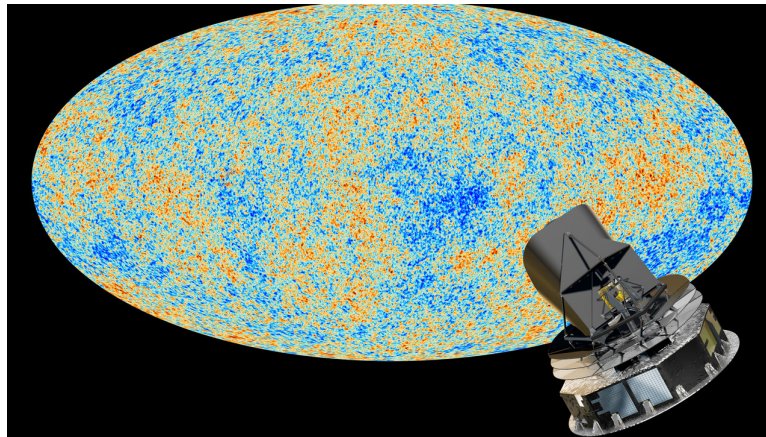
Ultra-High Energy
Cosmic Rays

Quiz: If SHDM is near the GUT scale, how do we probe it?



Galaxy-Size Collider!

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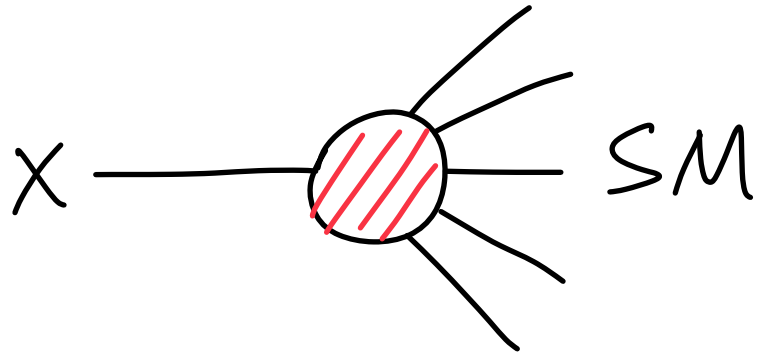
CMB Polarisation



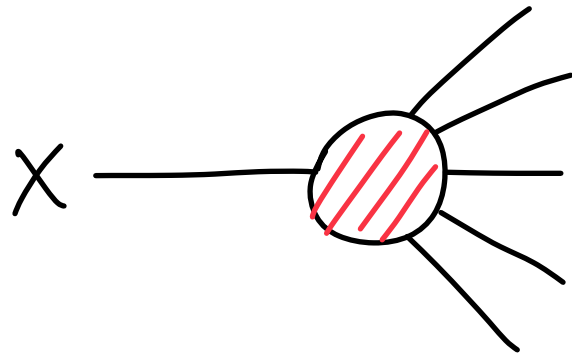
Ultra-High Energy
Cosmic Rays



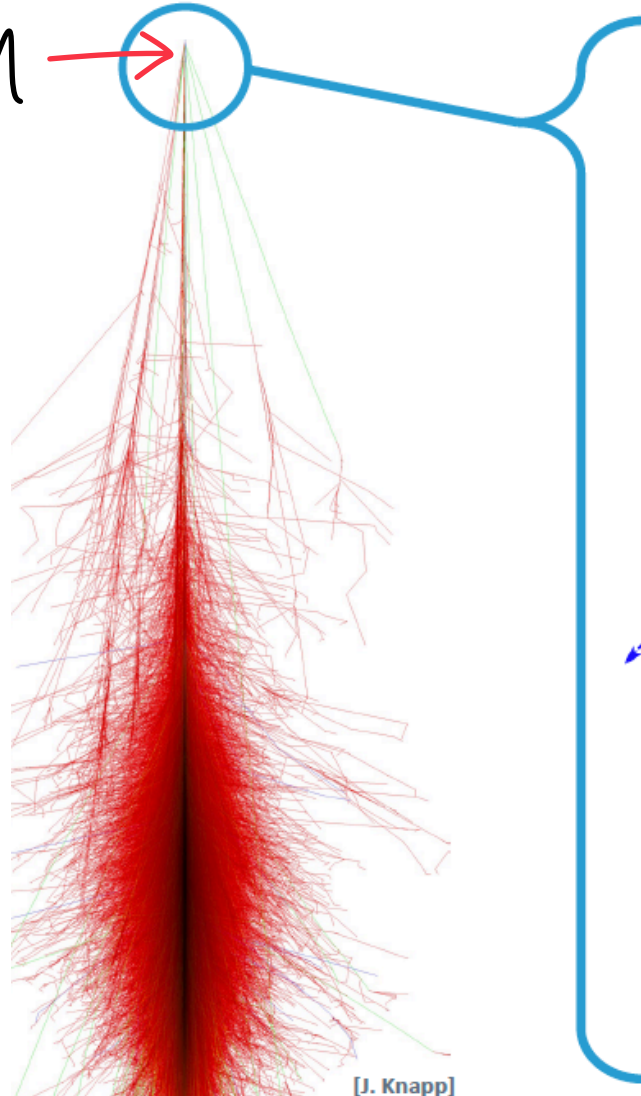
SHDM Decay Scenario: Perturbative, with a small coupling to the SM



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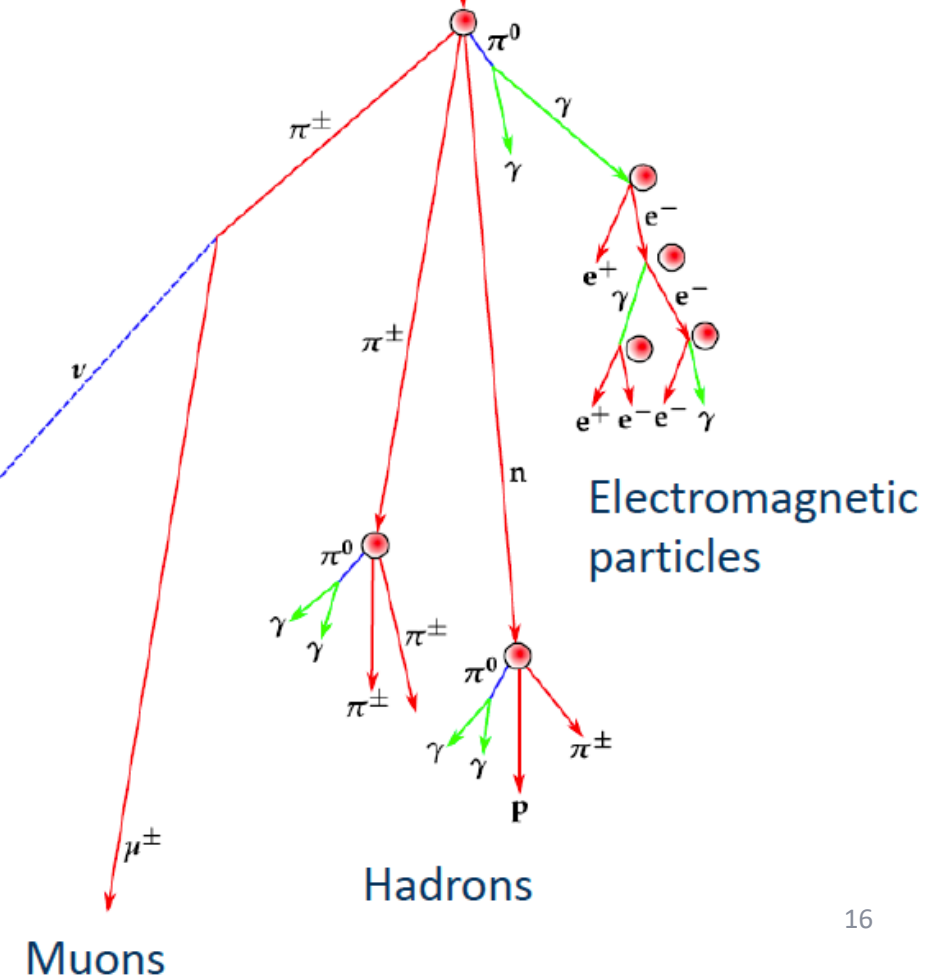
SM



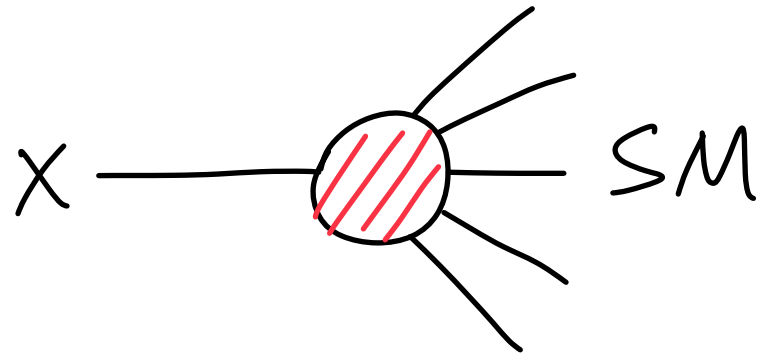
[J. Knapp]

Primary cosmic-ray particle

Nucleus from the Earth



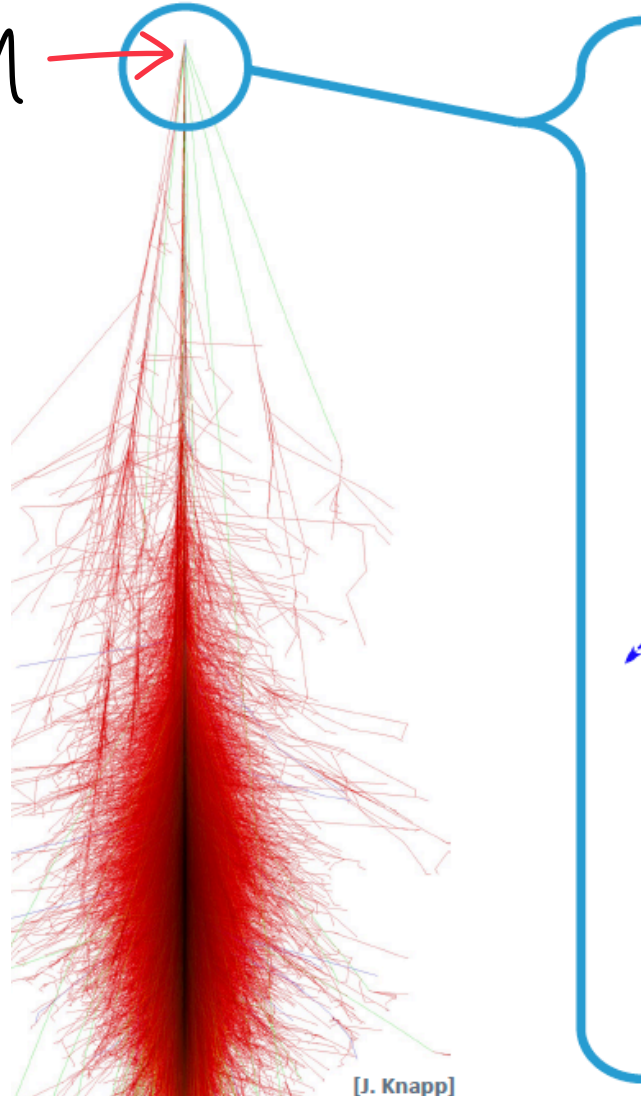
SHDM Decay Scenario: Perturbative, with a small coupling to the SM



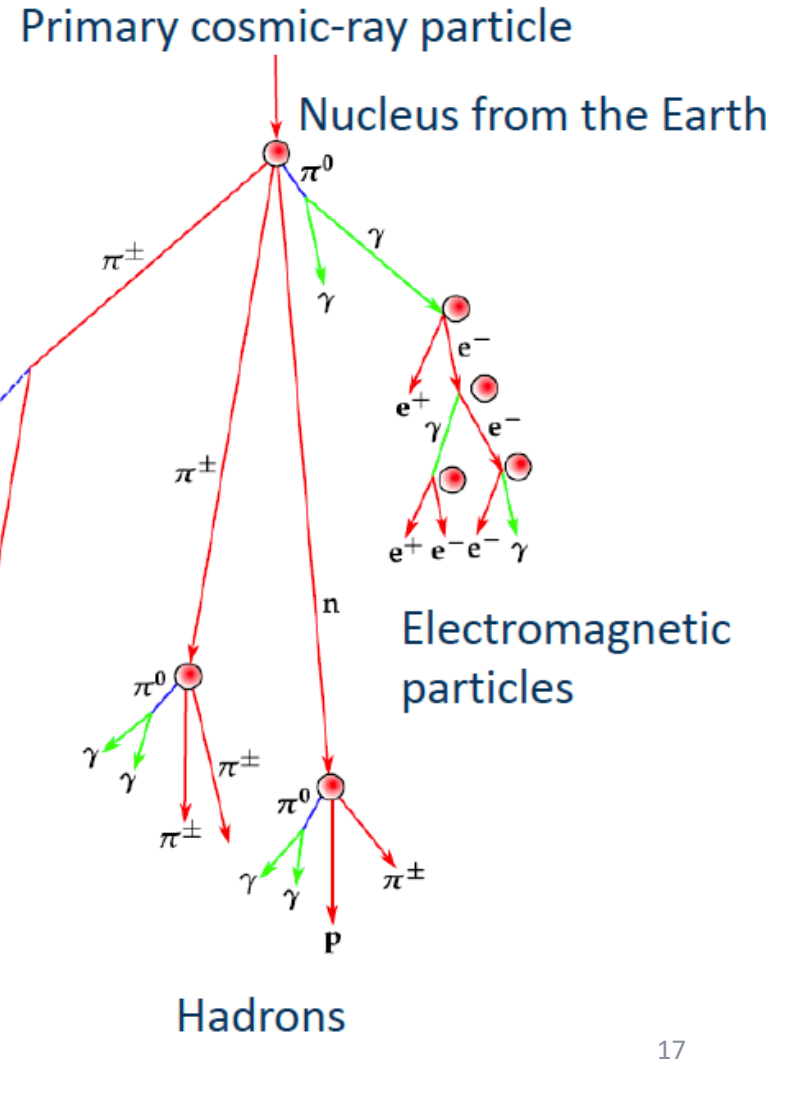
$$\mathcal{L} = \frac{g}{\Lambda^n} X [\text{SM}]$$

$$\tau_X \sim \frac{1}{g^2} \left(\frac{\Lambda}{M_X} \right)^{2n}$$

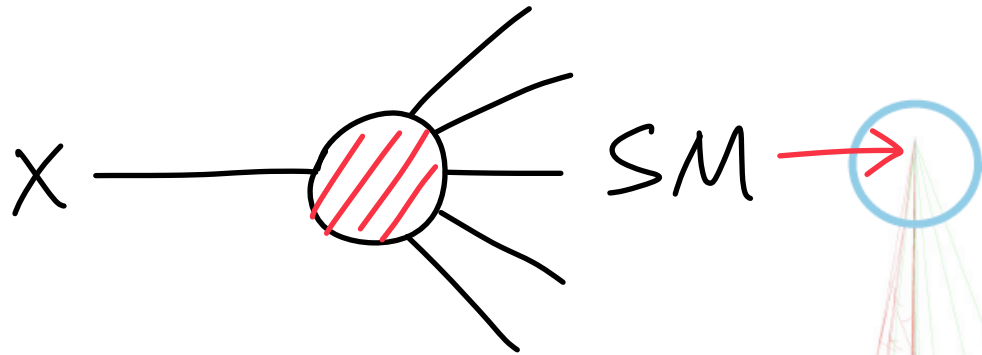
$$\Lambda \sim \text{GUT} \sim 10^{16} \text{ GeV}$$



[J. Knapp]



SHDM Decay Scenario: Perturbative, with a small coupling to the SM

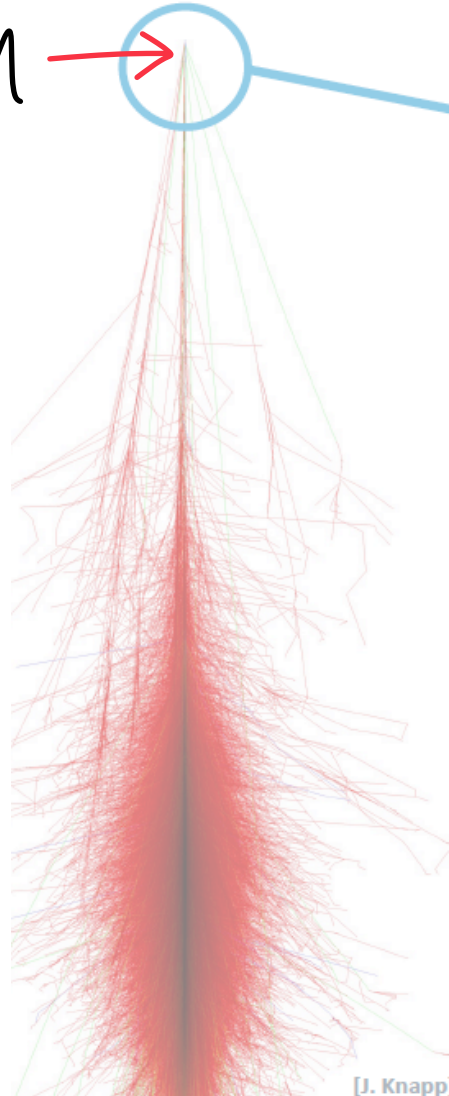


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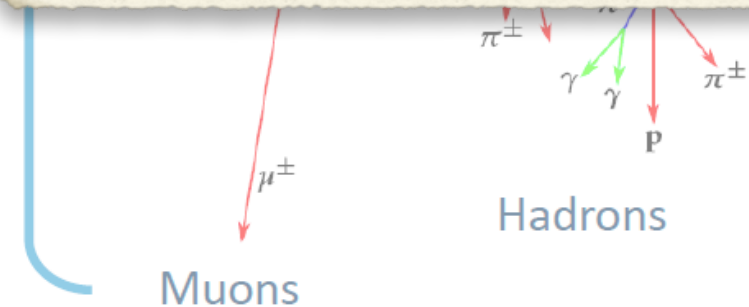
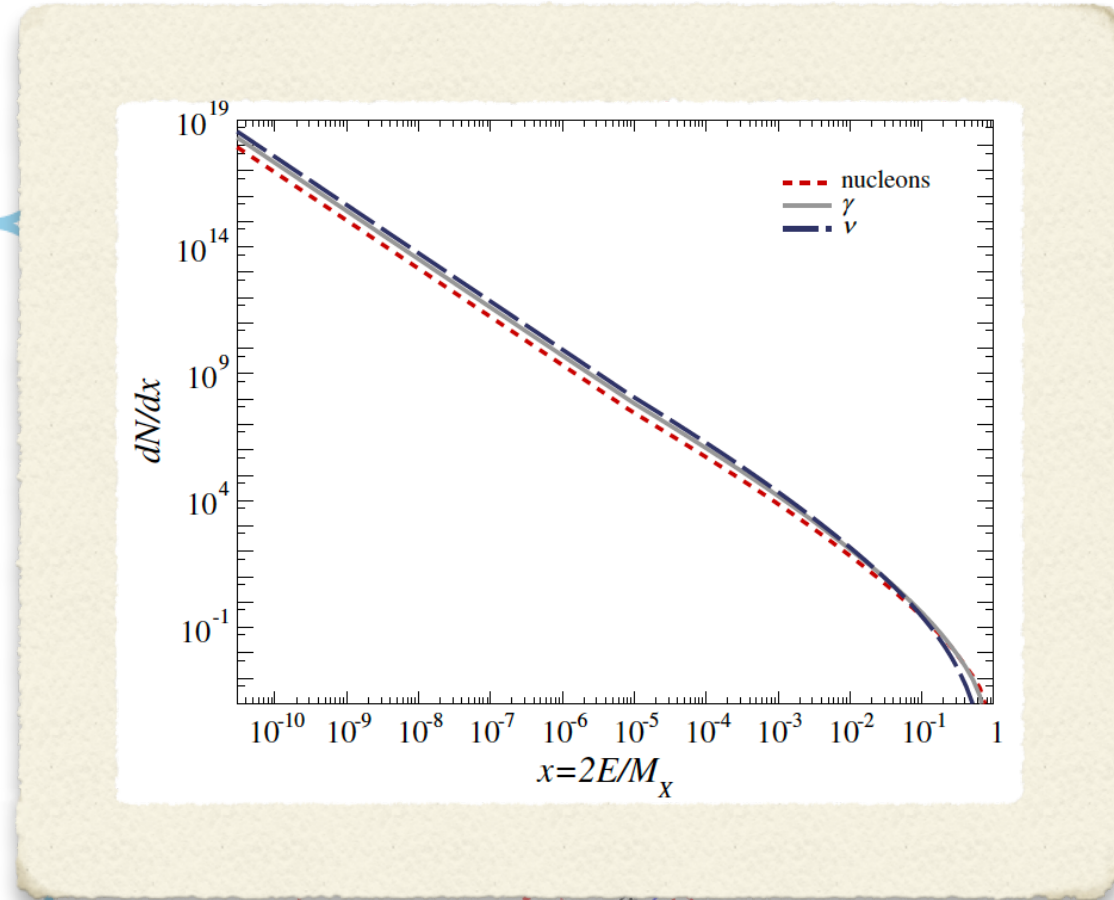
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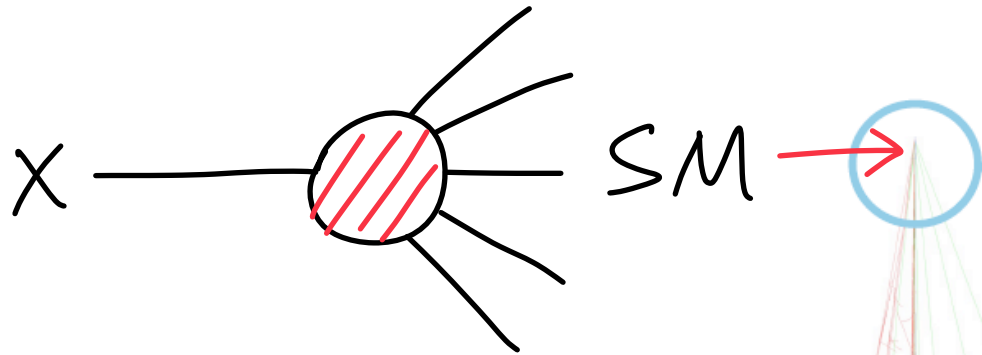
u



[J. Knapp]



SHDM Decay Scenario: Perturbative, with a small coupling to the SM



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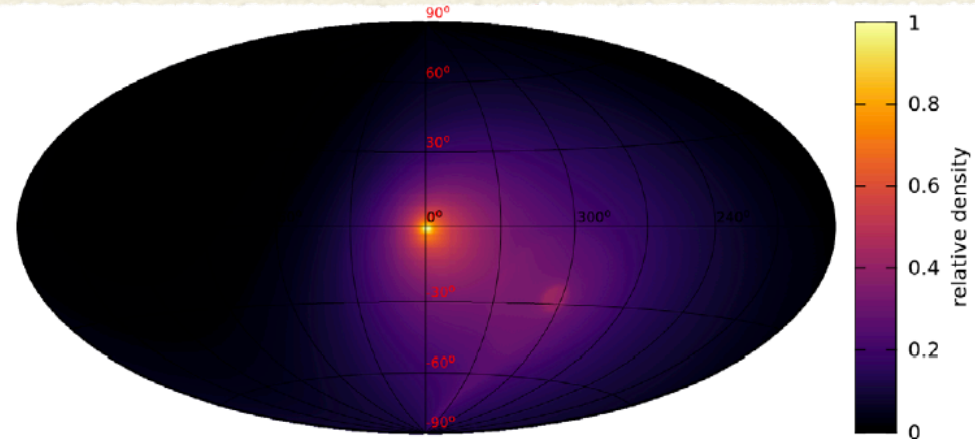
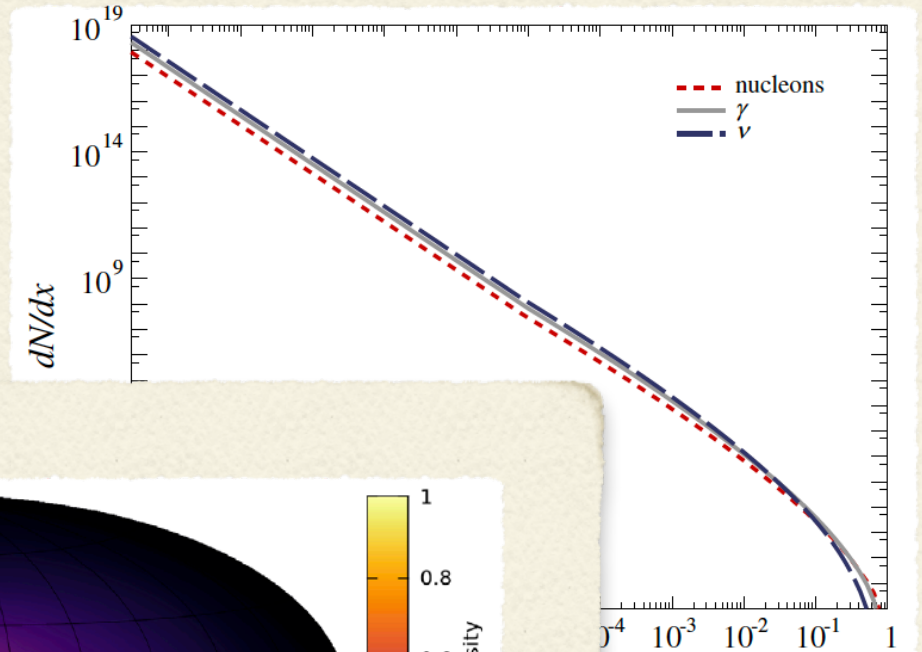
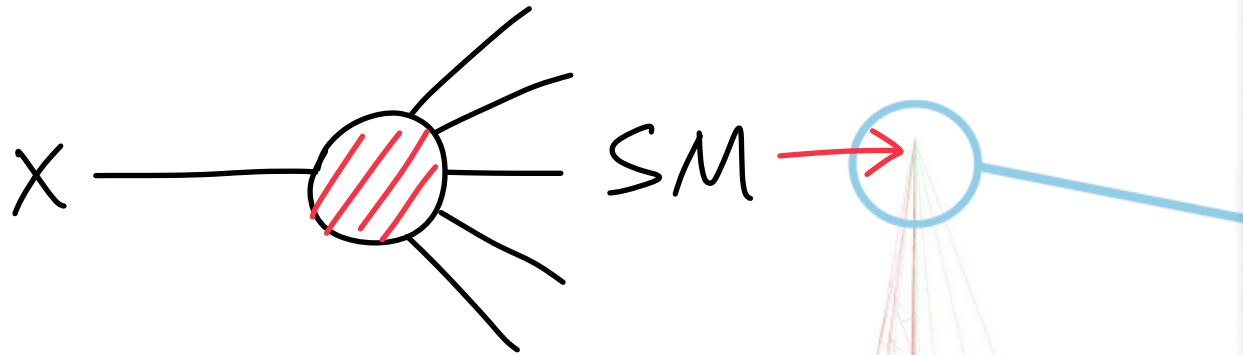


FIG. 2. Signal term of the directional density, $\delta\mu(\mathbf{n}, E = 32 \text{ EeV})$, as expected to be observed at the Pierre Auger Observatory in Galactic coordinates.

SHDM Decay Scenario: Perturbative, with a

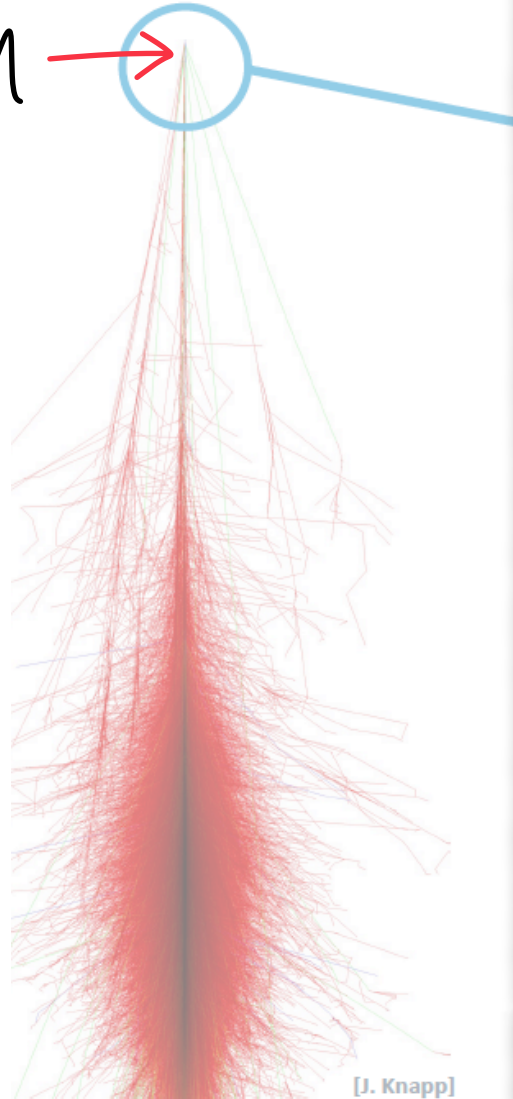


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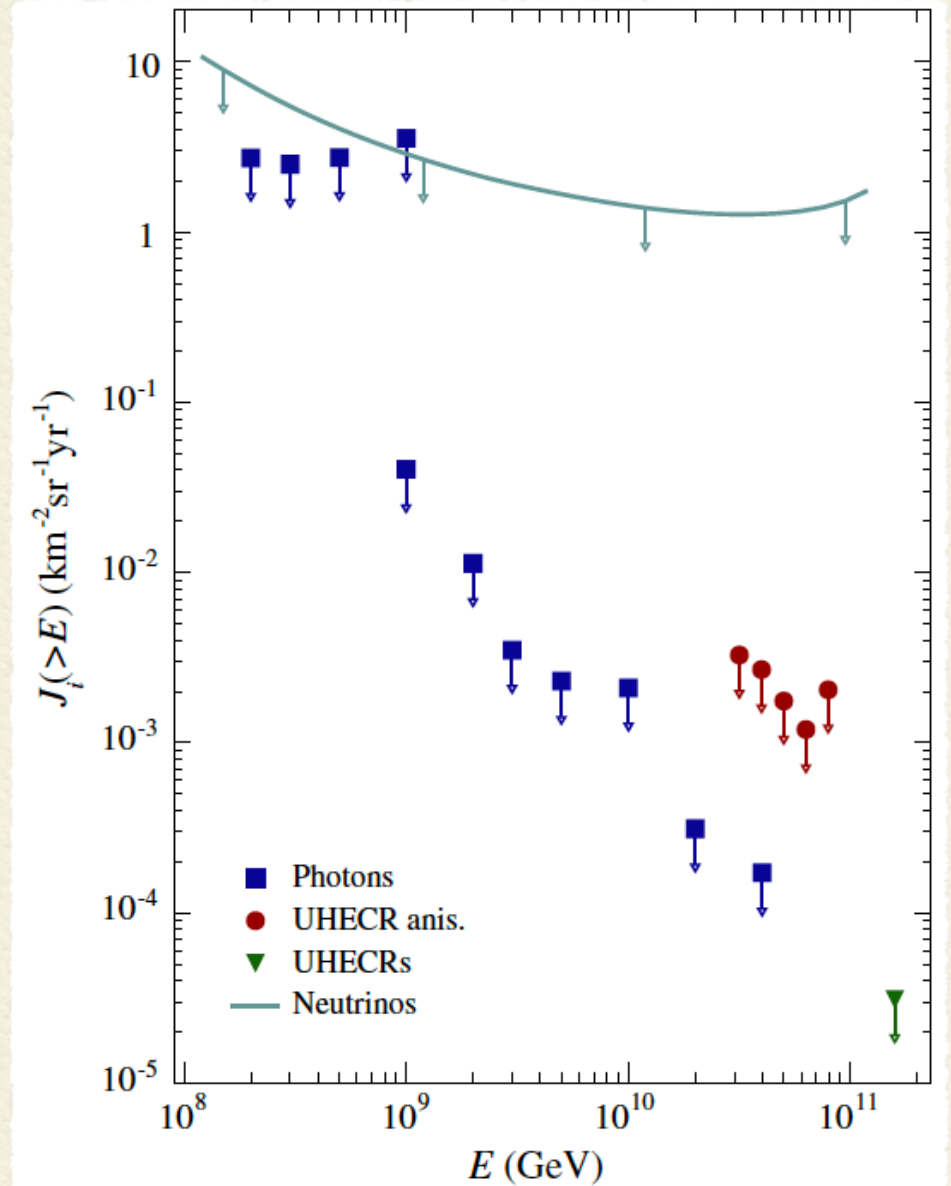
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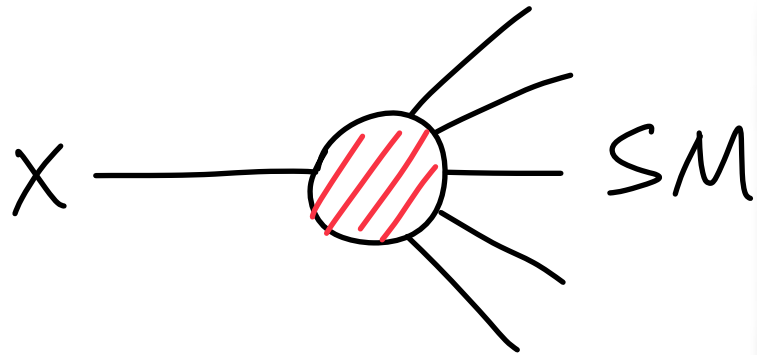
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[J. Knapp]



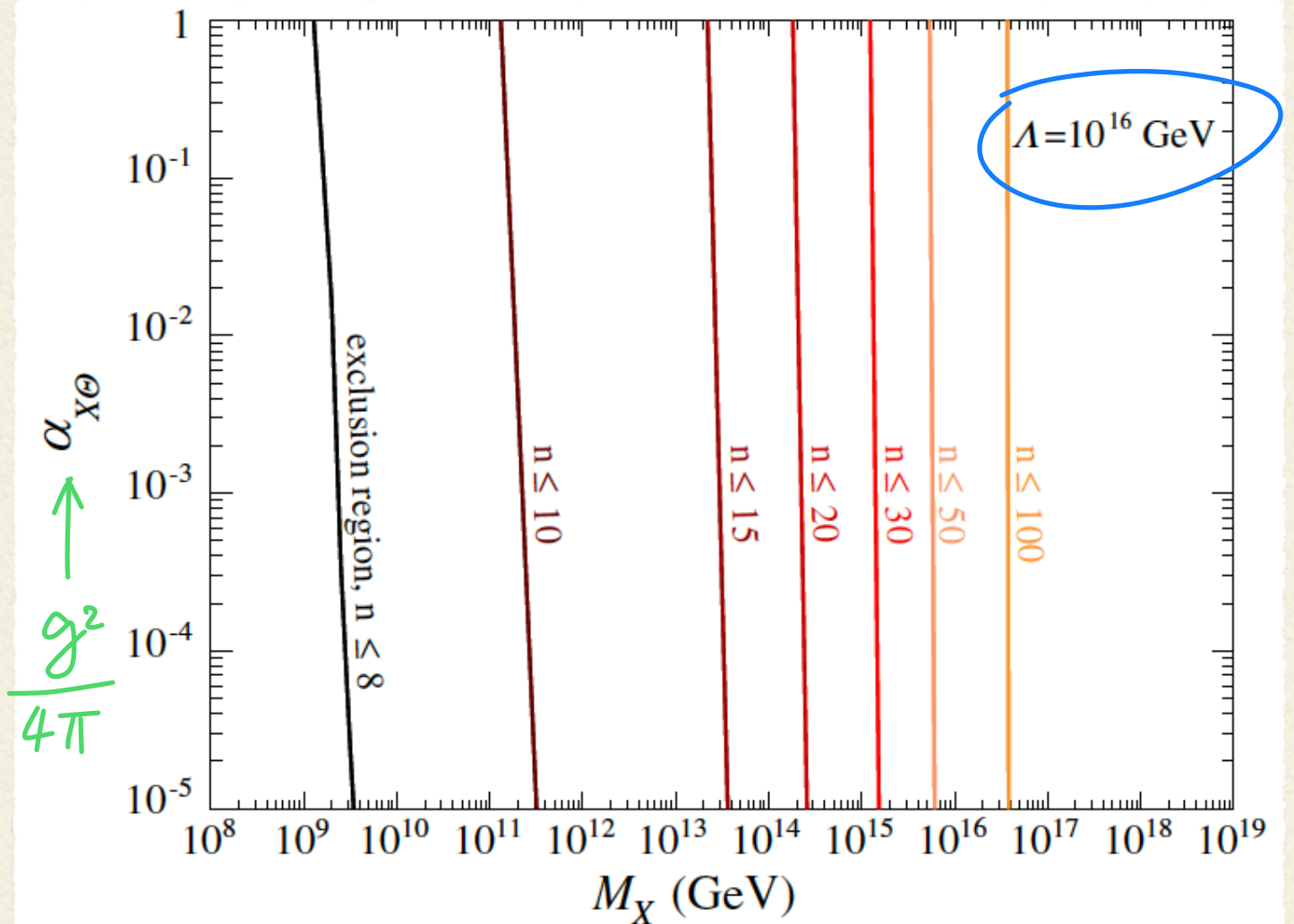
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Quiz: A physicist's Dark Matter Nightmare



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No WIMPs! No SUSY!



Quiz: A physicist's Dark Matter Nightmare



No WIMPs! No SUSY!



Quiz: A physicist's Dark Matter Nightmare



No WIMPs! No SUSY!

DM are not particles!



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$$g = 0$$



No WIMPs! No SUSY!

DM are not particles!

DM don't talk to the SM



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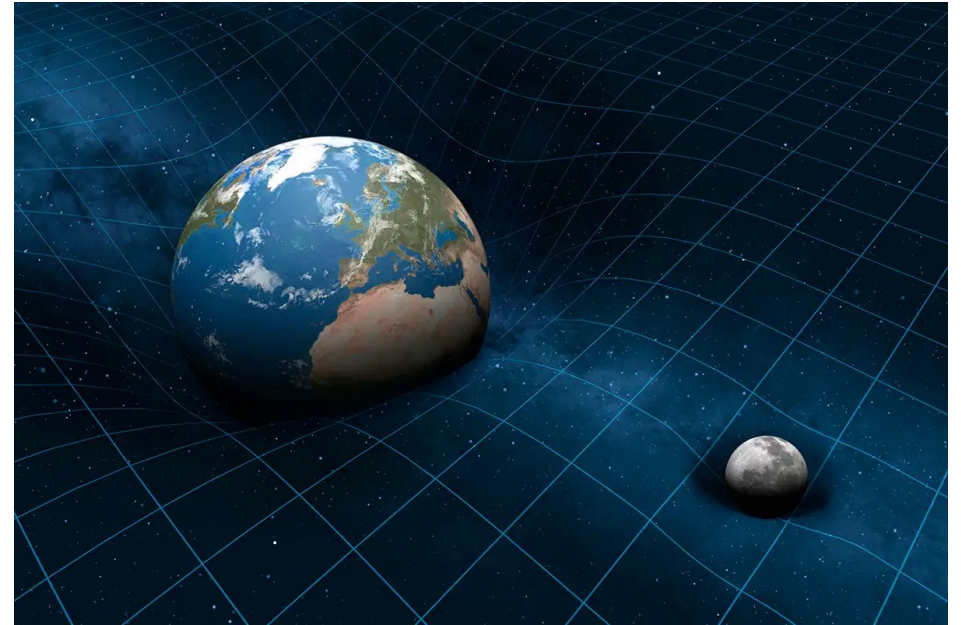
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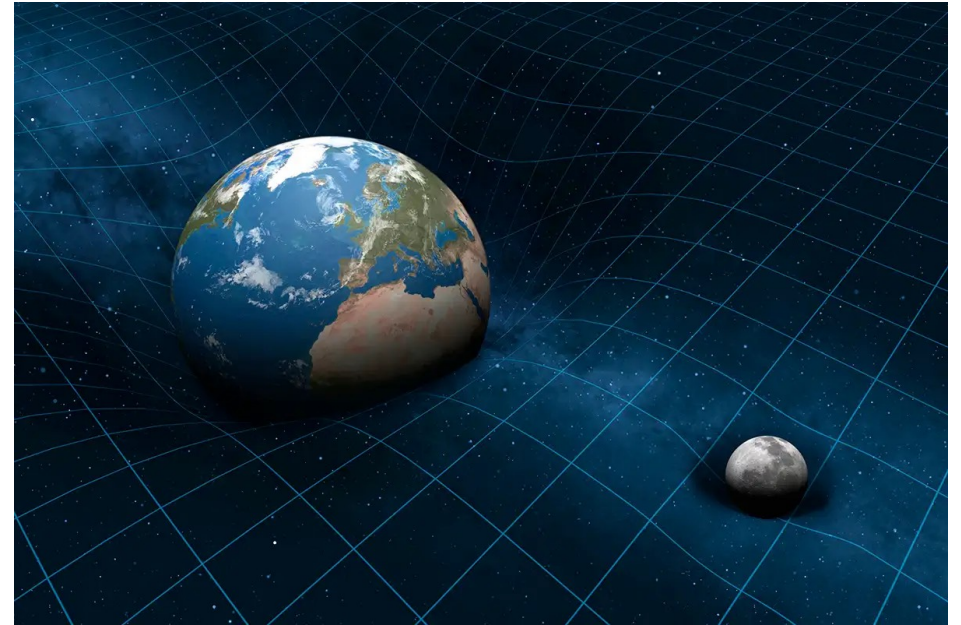
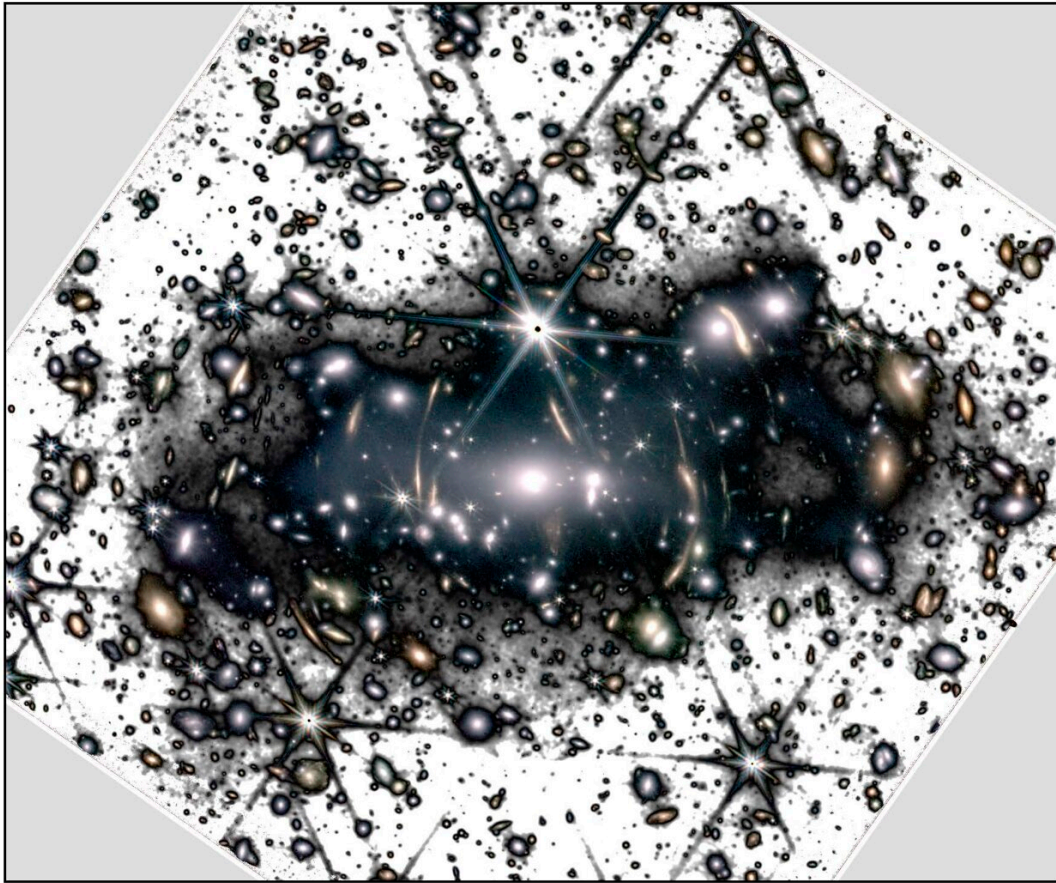
However..



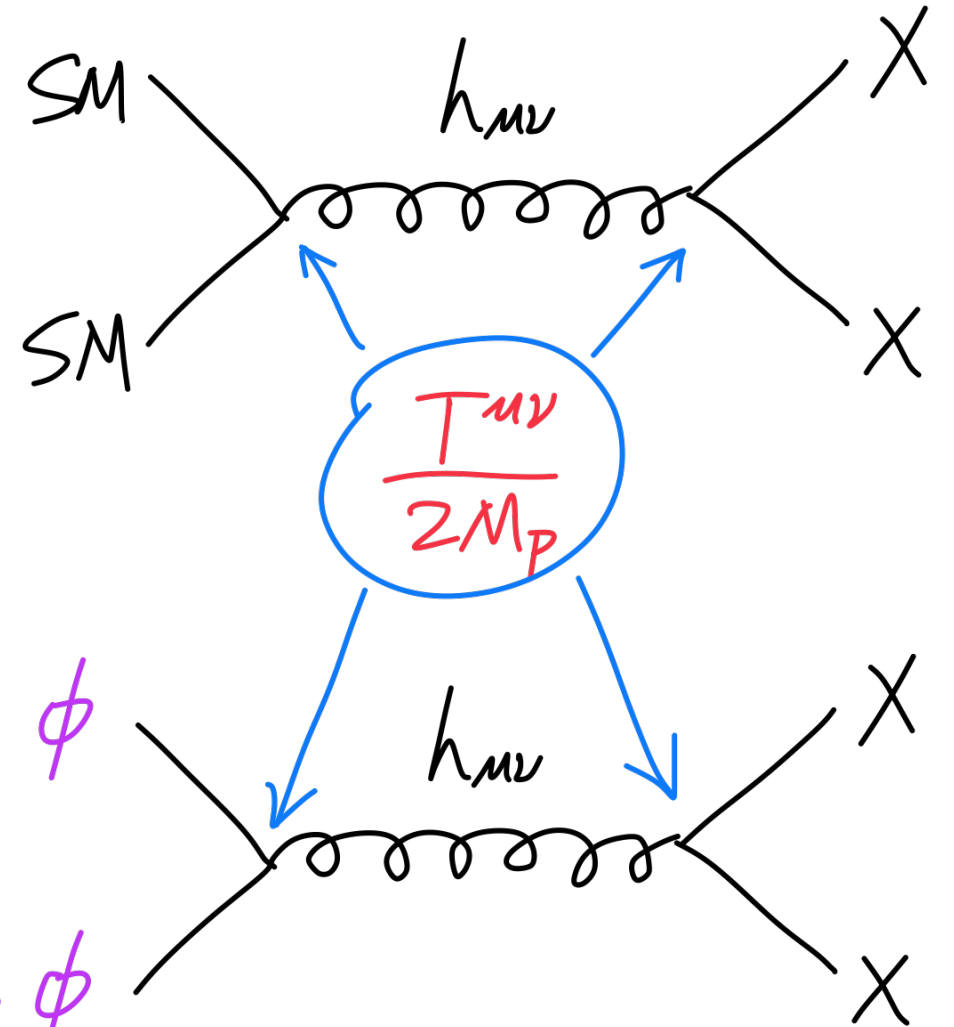
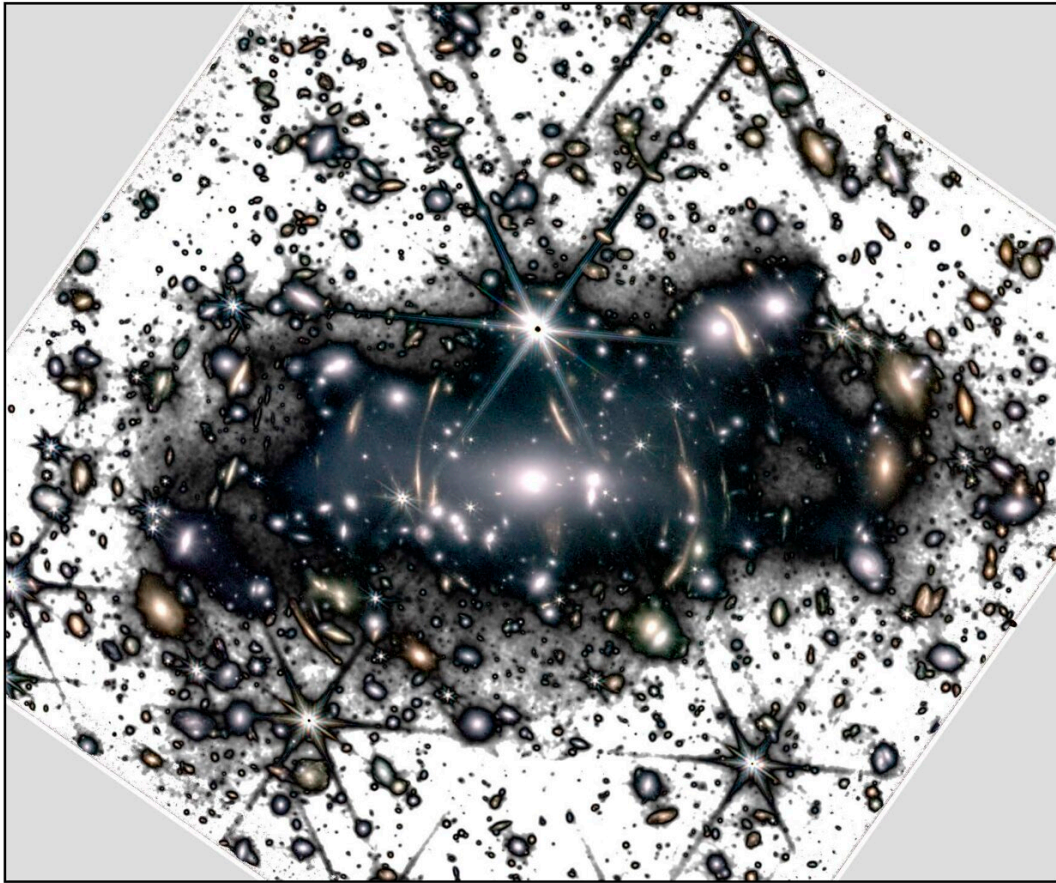
However..



Gravity connects ALL



Purely gravitational SHDM can be produced around reheating



Inflaton $\rightarrow \phi$

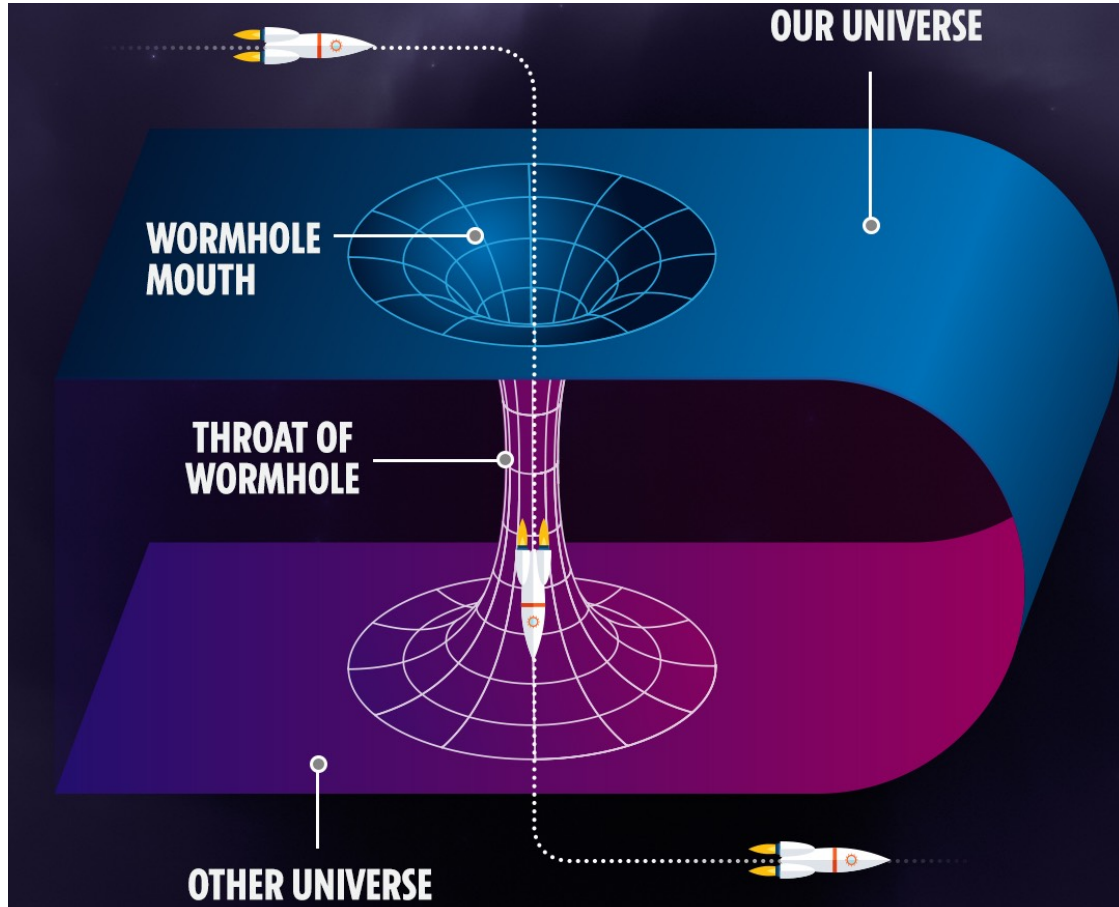
Gravity eats ALL



Gravity eats ALL

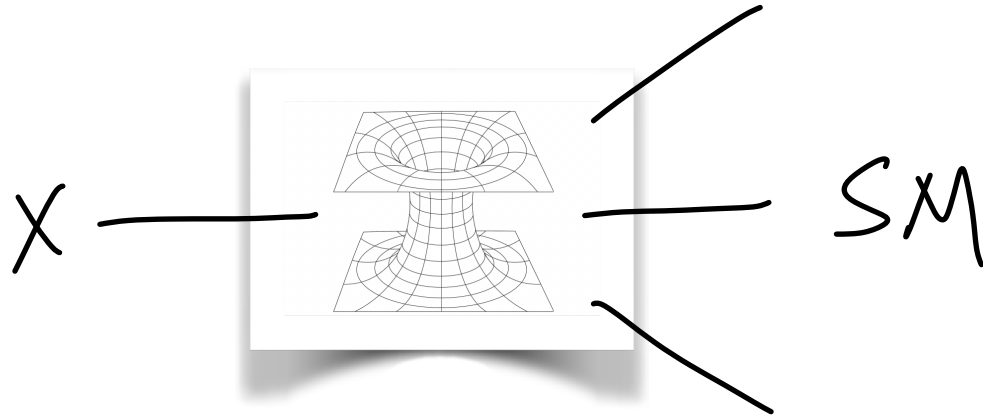
- Assume the DM is stable
- Protected by a global symmetry
- The new quantum number is respected by all particle interactions
- Even such a DM particle can decay
- Through gravitational instanton
- Quantum gravity effects!

Gravity eats ALL

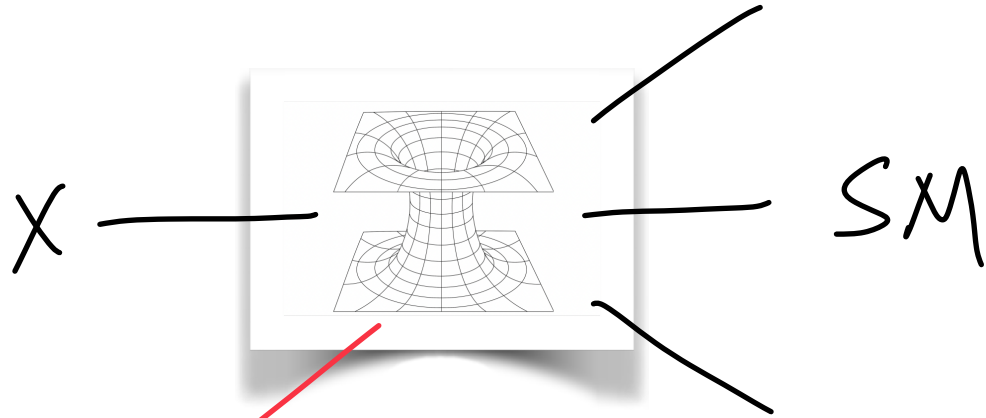


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SHDM Decay Scenario: Non-perturbative



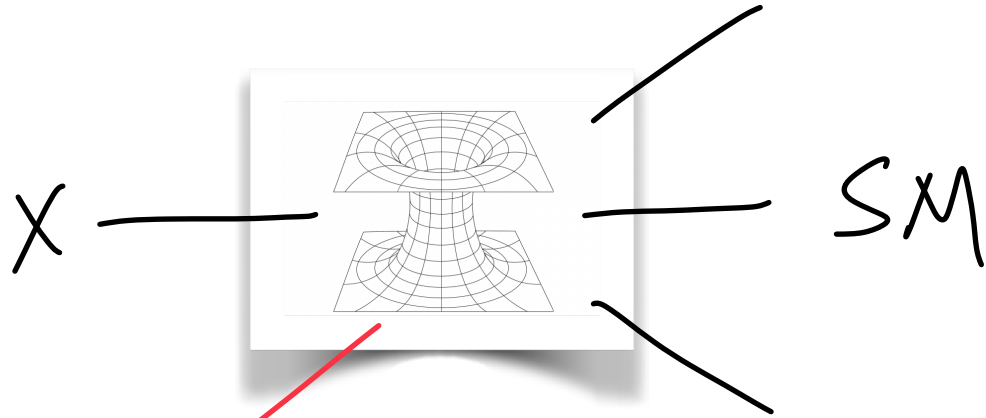
SHDM Decay Scenario: Non-perturbative



$$e^{-S}, \quad S \sim \left(\frac{R}{L_P}\right)^2$$

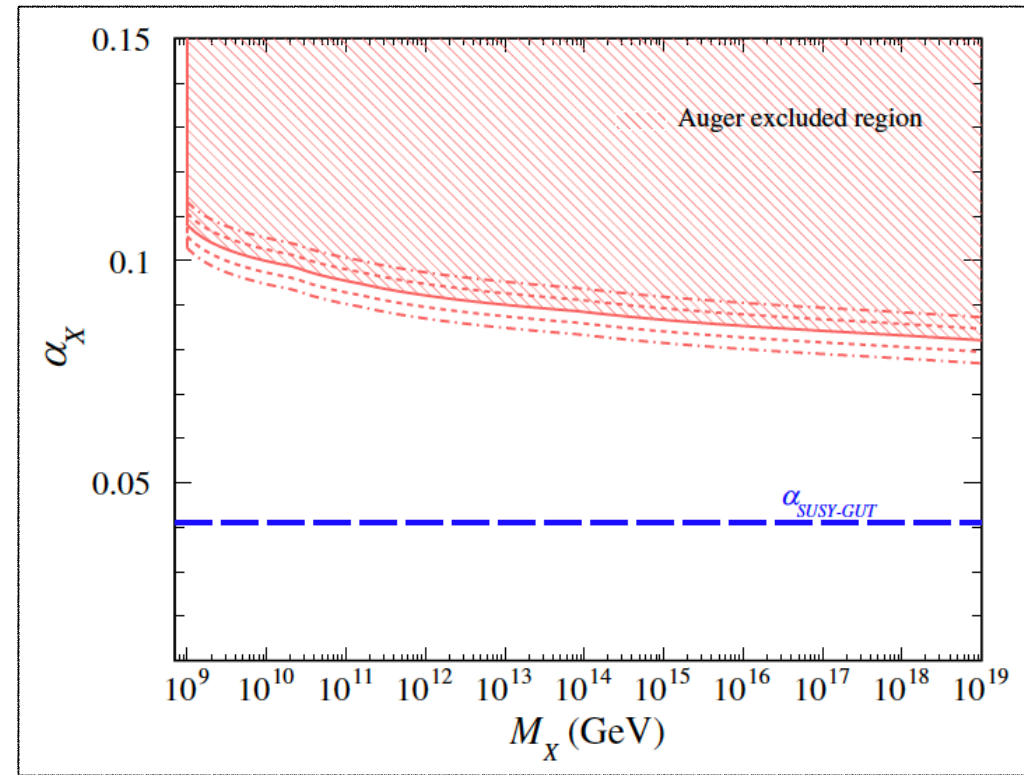
$$\tau_X \sim \frac{e^{2S}}{M_X} \sim \frac{e^{\frac{4\pi}{\alpha}}}{M_X}$$

SHDM Decay Scenario: Non-perturbative

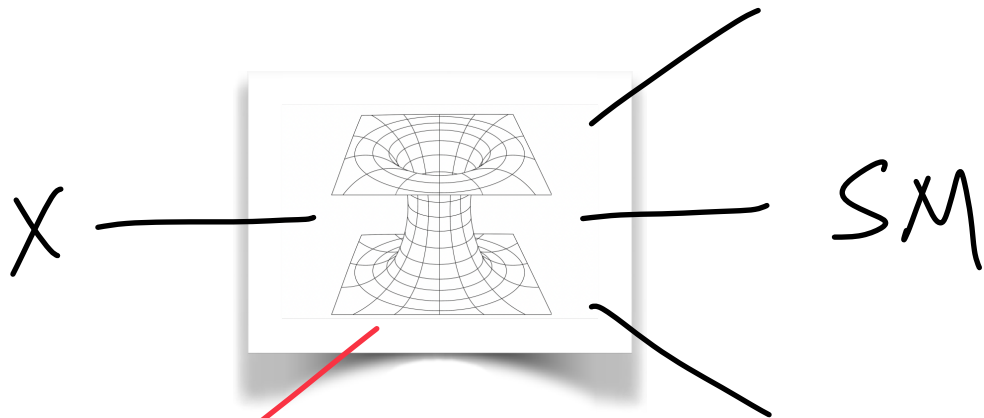


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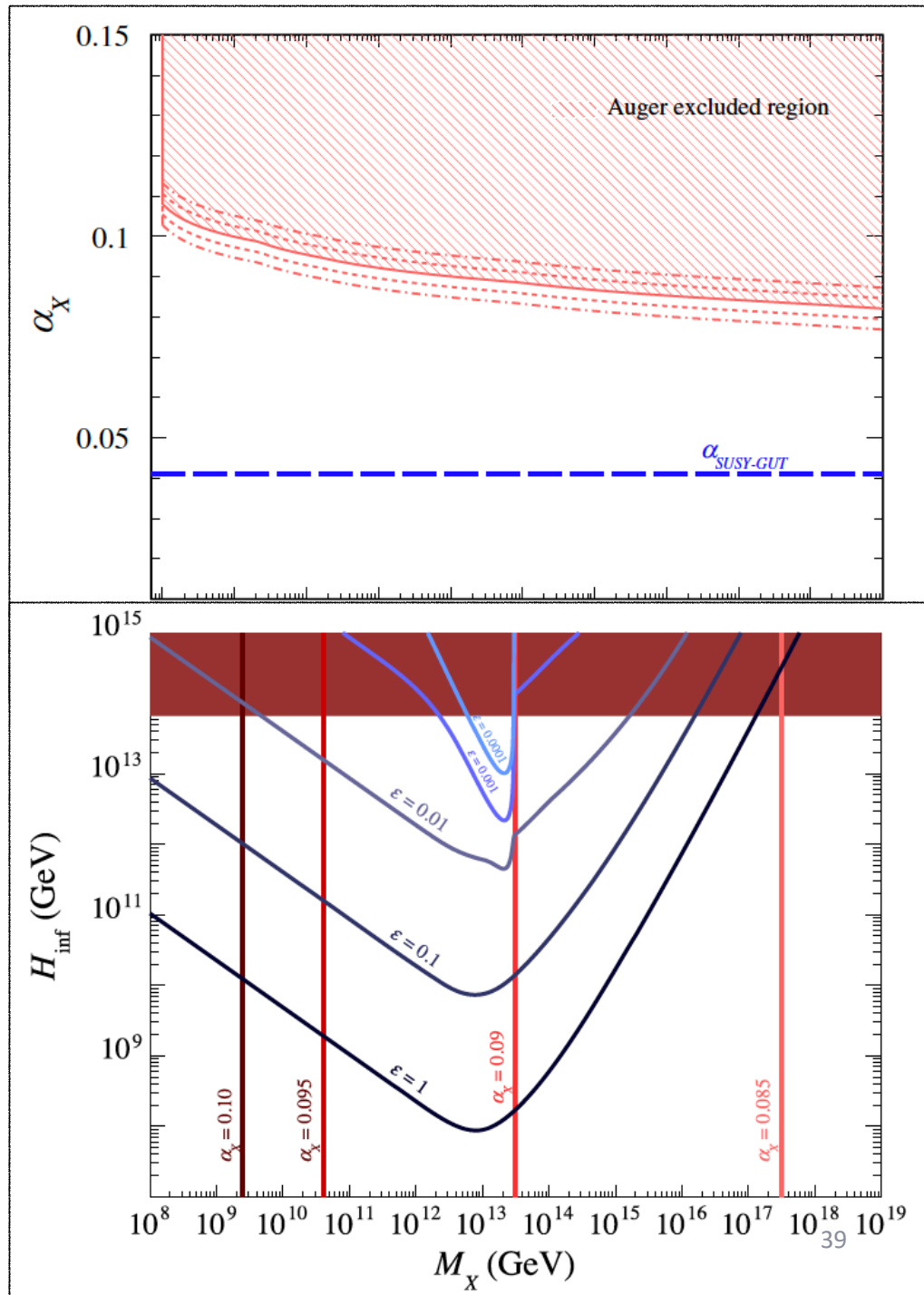


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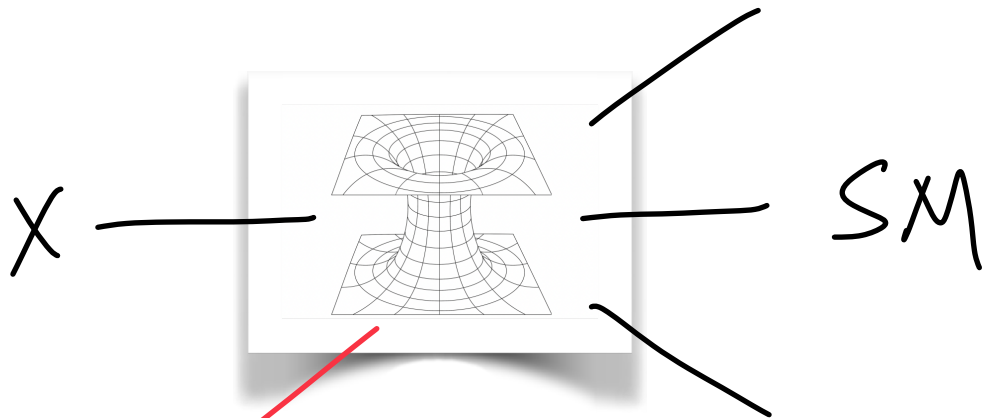


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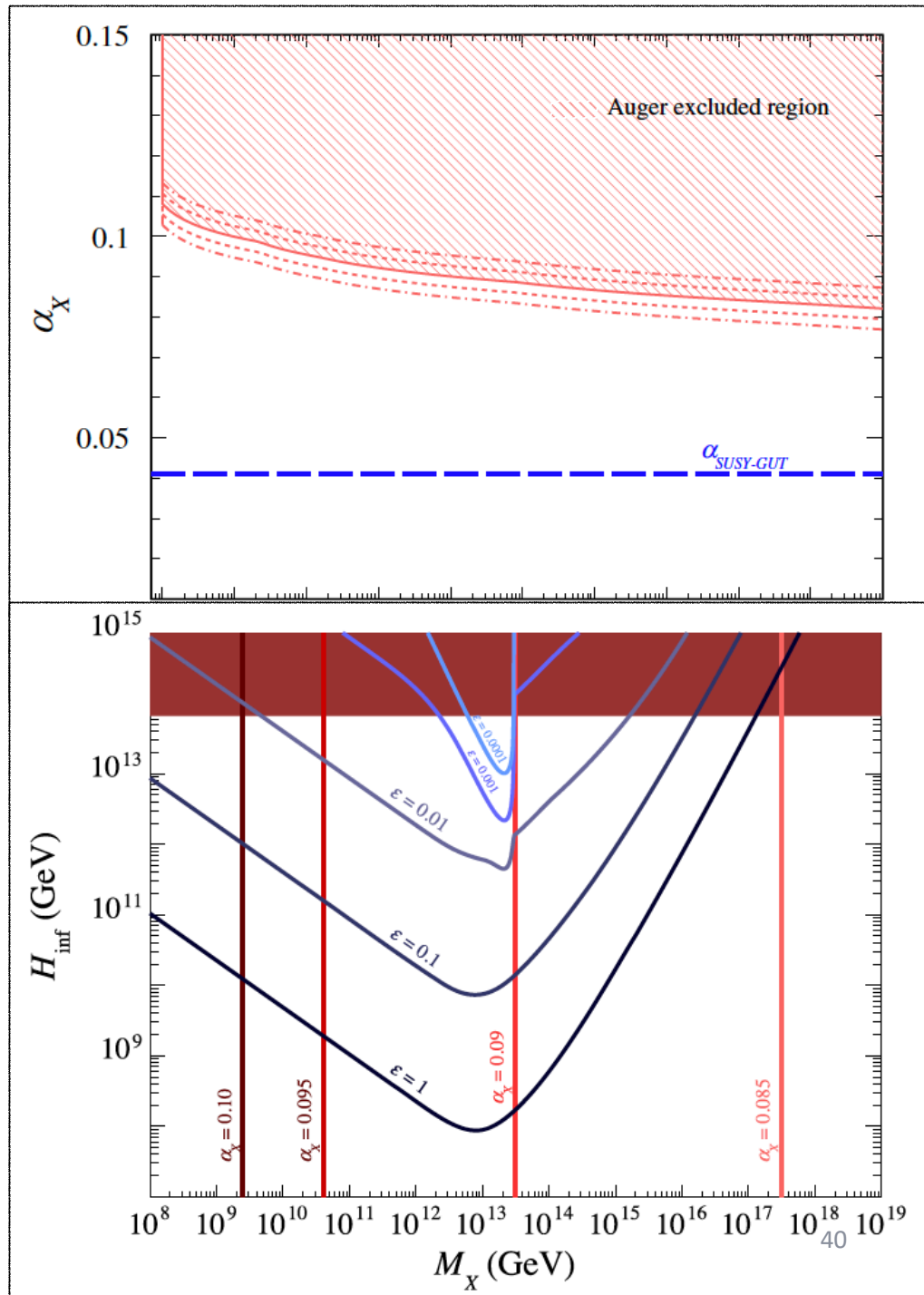


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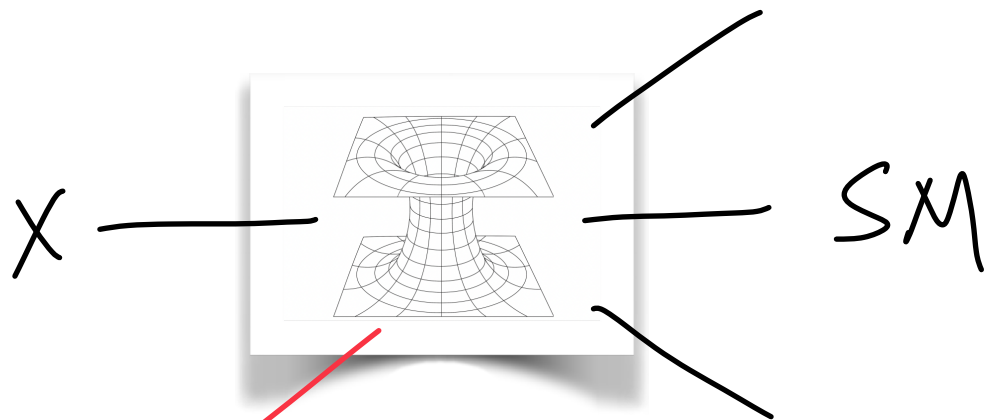


$$e^{-S}, \quad S \sim \left(\frac{R}{L_P}\right)^2 > 80$$

$$\tau_X \sim \frac{e^{2S}}{M_X} \sim \frac{e^{\frac{4\pi}{\alpha}}}{M_X} > 10^{22} \text{ y}$$



SHDM Decay Scenario: Non-perturbative



$$e^{-S}, \quad S \sim \left(\frac{R}{L_P}\right)^2 > 80 \approx \frac{8\pi^2}{g_s^2}$$

$$\tau_X \sim \frac{e^{2S}}{M_X} \sim \frac{e^{\frac{4\pi}{g_s^2}}}{M_X} > 10^{22} \text{ y}$$

