Indirect Constraints on Third Generation Baryon Number Violation

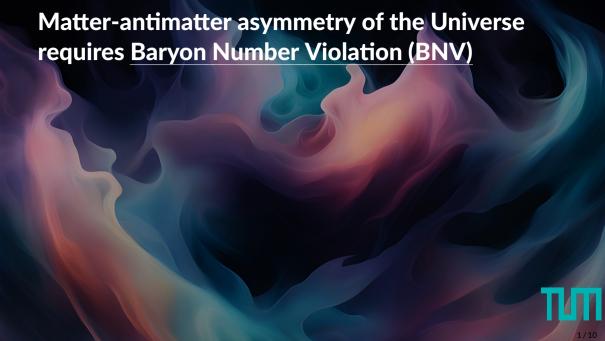
Gael Finauri

Quirks 2024

Zadar, Croatia - 18 June 2024

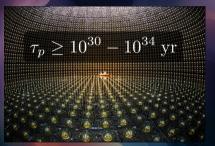


TUT



Matter-antimatter asymmetry of the Universe requires Baryon Number Violation (BNV)

Proton lifetime τ_p is the strongest constraint on BNV:

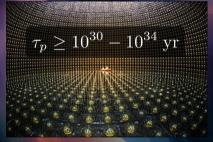


Super-Kamiokande, Japan



Matter-antimatter asymmetry of the Universe requires Baryon Number Violation (BNV)

Proton lifetime τ_p is the strongest constraint on BNV:



Super-Kamiokande, Japan

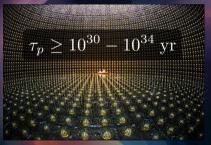
suggesting New Physics from very high scales

$$\Lambda_{\rm BNV} \sim 10^{15} - 10^{16} \, {\rm GeV}$$



Matter-antimatter asymmetry of the Universe requires Baryon Number Violation (BNV)

Proton lifetime τ_p is the strongest constraint on BNV:



Super-Kamiokande, Japan

suggesting New Physics from very high scales

$$\Lambda_{\rm BNV} \sim 10^{15} - 10^{16} \, {\rm GeV}$$

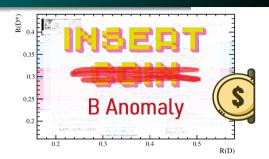
But what if BNV needs third generation quarks?

B-anomalies



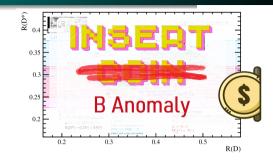
B-anomalies

explained by New Physics at scales $\Lambda_{\rm fl} \sim 1-10~{\rm TeV}$



B-anomalies

explained by New Physics at scales $\Lambda_{\rm fl} \sim 1-10~{\rm TeV}$



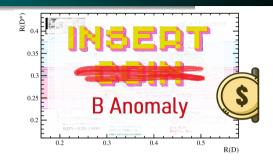
Natural questions:

• If BNV comes only from b quarks, can $\Lambda_{\rm BNV} \sim \Lambda_{\rm fl} \ll 10^{15}$ GeV be possible?



B-anomalies

explained by New Physics at scales $\Lambda_{\rm fl} \sim 1-10~{\rm TeV}$



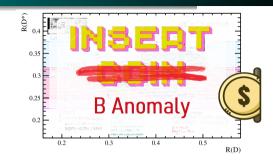
Natural questions:

- If BNV comes only from b quarks, can $\Lambda_{\rm BNV} \sim \Lambda_{\rm fl} \ll 10^{15}$ GeV be possible?
- Would it lead to observable BNV B decays?



B-anomalies

explained by New Physics at scales $\Lambda_{\rm fl} \sim 1-10~{\rm TeV}$



Natural questions:

- If BNV comes only from b quarks, can $\Lambda_{\mathsf{BNV}} \sim \Lambda_{\mathsf{fl}} \ll 10^{15}$ GeV be possible?
- Would it lead to observable BNV B decays?

Answer: Don't know!

Proton would still decay through virtual b quarks, constraining Λ_{BNV}



Assumption: BNV happens only if at least a b quark is involved (in WET).



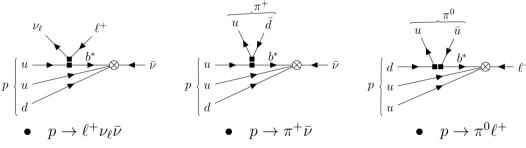
Assumption: BNV happens only if at least a b quark is involved (in WET). Goal: as $\Gamma \propto 1/\Lambda_{\rm RNV}^4$, find the smallest $\Lambda_{\rm BNV}$ allowed by proton decay

Assumption: BNV happens only if at least a b quark is involved (in WET).

Goal: as $\Gamma \propto 1/\Lambda_{\rm BNV}^4$, find the smallest $\Lambda_{\rm BNV}$ allowed by proton decay

 \Rightarrow estimate the <u>largest BNV</u> B decay rate

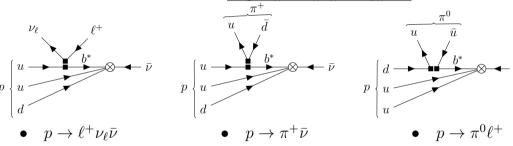
Assumption: BNV happens only if at least a b quark is involved (in WET). Goal: as $\Gamma \propto 1/\Lambda_{\rm BNV}^4$, find the smallest $\Lambda_{\rm BNV}$ allowed by proton decay \Rightarrow estimate the largest BNV B decay rate



mediated by weak interaction and four-fermion BNV operators (⊗) from SMEFT:



Assumption: BNV happens only if at least a b quark is involved (in WET). Goal: as $\Gamma \propto 1/\Lambda_{\rm BNV}^4$, find the smallest $\Lambda_{\rm BNV}$ allowed by proton decay \Rightarrow estimate the largest BNV B decay rate

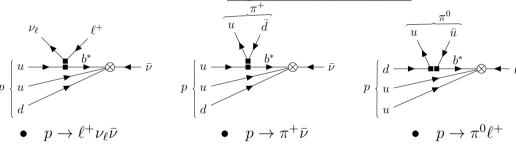


mediated by weak interaction and four-fermion BNV operators (⊗) from SMEFT:

 $\underline{\mathsf{most}\;\mathsf{suppressed}}\;\Gamma(p\to f)$



Assumption: BNV happens only if at least a b quark is involved (in WET). Goal: as $\Gamma \propto 1/\Lambda_{\rm BNV}^4$, find the smallest $\Lambda_{\rm BNV}$ allowed by proton decay \Rightarrow estimate the largest BNV B decay rate



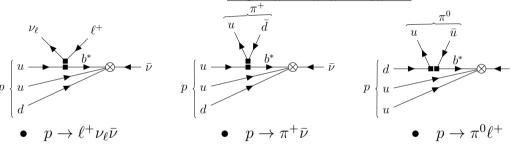
mediated by weak interaction and four-fermion BNV operators (\otimes) from SMEFT:



most suppressed $\Gamma(p \to f)$ smallest Λ_{BNV} from τ_p constraints



Assumption: BNV happens only if at least a b quark is involved (in WET). Goal: as $\Gamma \propto 1/\Lambda_{\rm BNV}^4$, find the smallest $\Lambda_{\rm BNV}$ allowed by proton decay \Rightarrow estimate the largest BNV B decay rate



mediated by weak interaction and four-fermion BNV operators (⊗) from SMEFT:



most suppressed $\Gamma(p \to f)$ smallest Λ_{BNV} from τ_p constraints

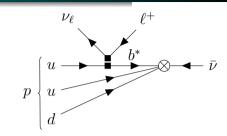


ightharpoonup highest $\Gamma(\bar{B} \to X\ell)$

Results for $p \to \ell^+ \nu_\ell \bar{\nu}$

Experimental bounds [Super-Kamiokande 2014]

- $\Gamma(p \to e^+ \nu \nu) < 1.23 \cdot 10^{-64} \text{ GeV}$
- $\Gamma(p \to \mu^+ \nu \nu) < 0.95 \cdot 10^{-64} \text{ GeV}$



Results for $p \to \ell^+ \nu_\ell \bar{\nu}$

Experimental bounds [Super-Kamiokande 2014]

- $\Gamma(p \to e^+ \nu \nu) < 1.23 \cdot 10^{-64} \text{ GeV}$
- $\Gamma(p \to \mu^+ \nu \nu) < 0.95 \cdot 10^{-64} \text{ GeV}$

 C_{ν} is the unknown dimensionless Wilson coefficient in the Weak Effective Theory (WET)

$$p \begin{cases} u & \ell^+ \\ u & \bar{\nu} \end{cases}$$

$$\Gamma(p \to \ell^+ \nu_\ell \bar{\nu}) = \frac{|C_\nu|^2}{\Lambda_{\mathsf{BNV}}^4} \frac{|V_{ub}|^2 G_F^2 m_p^7}{7680 \pi^3 m_b^2} 10^{-3} \mathsf{GeV}^4 \times \begin{cases} 1.028 \,, & \text{for } p \to e^+ \nu_e \bar{\nu} \\ 0.933 \,, & \text{for } p \to \mu^+ \nu_\mu \bar{\nu} \end{cases}$$



Results for $p \to \ell^+ \nu_\ell \bar{\nu}$

Experimental bounds [Super-Kamiokande 2014]

- $\Gamma(p \to e^+ \nu \nu) < 1.23 \cdot 10^{-64} \text{ GeV}$
- $\Gamma(p \to \mu^+ \nu \nu) < 0.95 \cdot 10^{-64} \text{ GeV}$

 C_{ν} is the unknown dimensionless Wilson coefficient in the Weak Effective Theory (WET)

$$p \begin{cases} u & b^* \\ u & \bar{\nu} \end{cases}$$

$$\Gamma(p \to \ell^+ \nu_\ell \bar{\nu}) = \frac{|C_\nu|^2}{\Lambda_{\rm RNV}^4} \frac{|V_{ub}|^2 G_F^2 m_p^7}{7680 \pi^3 m_b^2} 10^{-3} {\rm GeV}^4 \times \begin{cases} 1.028 \,, & \text{for } p \to e^+ \nu_e \bar{\nu} \\ 0.933 \,, & \text{for } p \to \mu^+ \nu_u \bar{\nu} \end{cases}$$

Using the experimental constraints

$$\left. \frac{\Lambda_{\rm BNV}}{\sqrt{|C_\nu|}} \right|_{p \to e^+\nu_e \bar{\nu}} > 6.59 \cdot 10^9 \; {\rm GeV} \qquad \qquad \frac{\Lambda_{\rm BNV}}{\sqrt{|C_\nu|}} \bigg|_{p \to \mu^+\nu_\mu \bar{\nu}} > 6.86 \cdot 10^9 \; {\rm GeV} \right.$$

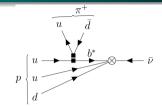
rather **high!** Already showing that $\Lambda_{\text{BNV}} \sim \Lambda_{\text{fl}}$ is ruled out!

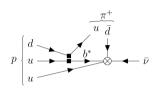


Experimental bounds [Super-Kamiokande 2014]

$$\Gamma(p\to\pi^+\nu)<5.35\cdot 10^{-65}~{\rm GeV}$$

Assume factorization, true up to $\mathcal{O}(1)$ corrections

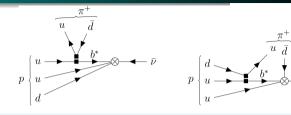




Experimental bounds [Super-Kamiokande 2014]

$$\Gamma(p\to\pi^+\nu)<5.35\cdot 10^{-65}~{\rm GeV}$$

Assume factorization, true up to $\mathcal{O}(1)$ corrections

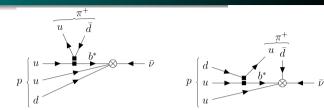


$$\Gamma(p \to \pi^+ \bar{\nu}) = \frac{|C_{\nu}|^2}{\Lambda_{\text{PNN}}^4} \frac{|V_{ud}|^2 |V_{ub}|^2 G_F^2 m_p^5 f_{\pi}^2}{1024\pi m_{\nu}^2} (2.18 \cdot 10^{-5} \text{GeV}^4)$$



Experimental bounds [Super-Kamiokande 2014] $\Gamma(p \to \pi^+
u) < 5.35 \cdot 10^{-65} \, {\rm GeV}$

Assume factorization, true up to $\mathcal{O}(1)$ corrections



$$\Gamma(p \to \pi^+ \bar{\nu}) = \frac{|C_{\nu}|^2}{\Lambda_{\text{PNN}}^4} \frac{|V_{ud}|^2 |V_{ub}|^2 G_F^2 m_p^5 f_{\pi}^2}{1024\pi m_{\pi}^2} (2.18 \cdot 10^{-5} \text{GeV}^4)$$

neglected suppressed contributions:

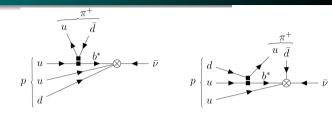
$$p\begin{bmatrix} d & & & \bar{v} \\ u & & & \bar{v} \\ u & & & & \bar{d} \end{bmatrix}\pi^+ \qquad p\begin{bmatrix} d & & & \bar{v} \\ u & & & \bar{d} \\ u & & & & \bar{d} \end{bmatrix}\pi^+$$





Experimental bounds [Super-Kamiokande 2014] $\Gamma(p \to \pi^+ \nu) < 5.35 \cdot 10^{-65} \, {\rm GeV}$

Assume factorization, true up to $\mathcal{O}(1)$ corrections



$$\Gamma(p \to \pi^+ \bar{\nu}) = \frac{|C_{\nu}|^2}{\Lambda_{\text{App}, \nu}^4} \frac{|V_{ud}|^2 |V_{ub}|^2 G_F^2 m_p^5 f_{\pi}^2}{1024\pi m_{\tau}^2} (2.18 \cdot 10^{-5} \text{GeV}^4)$$

neglected suppressed contributions:

Using the experimental constraints

$$\left. rac{\Lambda_{
m BNV}}{\sqrt{|C_
u|}}
ight|_{p
ightarrow \pi^+ ar
u} > 3.34 \cdot 10^9 \ {
m GeV}$$

5/10

less effective than leptonic decay

Results for $p \to \pi^0 \ell^+$

Strongest experimental bounds [Super-Kamiokande 2020]

•
$$\Gamma(p \to \pi^0 e^+) < 0.87 \cdot 10^{-66} \text{ GeV}$$

•
$$\Gamma(p \to \pi^0 \mu^+) < 1.30 \cdot 10^{-66} \text{ GeV}$$

two independent operators contributing



Results for $p \to \pi^0 \ell^+$

Strongest experimental bounds [Super-Kamiokande 2020]

$$\Gamma(p \to \pi^0 e^+) < 0.87 \cdot 10^{-66} \text{ GeV}$$

•
$$\Gamma(p \to \pi^0 \mu^+) < 1.30 \cdot 10^{-66} \text{ GeV}$$

two independent operators contributing

$$\begin{split} & \left. \Lambda_{\text{BNV}} \right|_{p \to \pi^0 e^+} > 6.23 \cdot 10^{10} \; \text{GeV} \left(|C_R^e|^2 + 0.0014 \text{Re} [C_L^{e\,*} C_R^e] + 0.304 |C_L^e|^2 \right)^{1/4}, \\ & \left. \Lambda_{\text{BNV}} \right|_{p \to \pi^0 \mu^+} > 5.63 \cdot 10^{10} \; \text{GeV} \left(|C_R^\mu|^2 + 0.283 \text{Re} [C_L^{\mu *} C_R^\mu] + 0.308 |C_L^\mu|^2 \right)^{1/4}, \end{split}$$

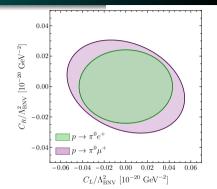
Results for $p \to \pi^0 \ell^+$

Strongest experimental bounds [Super-Kamiokande 2020]

- $\Gamma(p \to \pi^0 e^+) < 0.87 \cdot 10^{-66} \text{ GeV}$
- $\Gamma(p \to \pi^0 \mu^+) < 1.30 \cdot 10^{-66} \text{ GeV}$

two independent operators contributing

D constraints!



$$\begin{split} & \left. \Lambda_{\mathsf{BNV}} \right|_{p \to \pi^0 e^+} > 6.23 \cdot 10^{10} \; \mathsf{GeV} \left(|C_R^e|^2 + 0.0014 \mathsf{Re} [C_L^{e\,*} C_R^e] + 0.304 |C_L^e|^2 \right)^{1/4}, \\ & \left. \Lambda_{\mathsf{BNV}} \right|_{p \to \pi^0 \mu^+} > 5.63 \cdot 10^{10} \; \mathsf{GeV} \left(|C_R^\mu|^2 + 0.283 \mathsf{Re} [C_L^{\mu *} C_R^\mu] + 0.308 |C_L^\mu|^2 \right)^{1/4}, \end{split}$$

these coefficients are 10 times more constrained with respect to C_{ν}

We can estimate the branching ratio using $\Lambda_{\text{BNV}} > 6 \cdot 10^9 \text{ GeV}$

$$\mathcal{B}(ar{B} o X\ell) pprox rac{m_b^5}{2^{10} 3 \pi^3 \Gamma_{
m tot}^B \Lambda_{
m BNV}^4}$$

We can estimate the branching ratio using $\Lambda_{\text{BNV}} > 6 \cdot 10^9 \text{ GeV}$

$$\mathcal{B}(\bar{B} \to X\ell) \approx \frac{m_b^5}{2^{10}3\pi^3\Gamma_{\rm tot}^B\Lambda_{\rm RNV}^4} \approx (8|V_{cb}|G_F\Lambda_{\rm BNV}^2)^{-2}$$

We can estimate the branching ratio using $\Lambda_{\text{BNV}} > 6 \cdot 10^9 \text{ GeV}$

$$\mathcal{B}(\bar{B} \to X\ell) \approx \frac{m_b^5}{2^{10}3\pi^3\Gamma_{tot}^B\Lambda_{PNN}^4} \approx (8|V_{cb}|G_F\Lambda_{BNV}^2)^{-2} \lesssim \mathcal{O}(5\cdot 10^{-29})$$

showing that direct observation is ruled out: $\bar{B} \rightarrow X \ell$



We can estimate the branching ratio using $\Lambda_{\text{BNV}} > 6 \cdot 10^9 \text{ GeV}$

$$\mathcal{B}(\bar{B} \to X\ell) \approx \frac{m_b^5}{2^{10} 3\pi^3 \Gamma_{tot}^B \Lambda_{PNN}^4} \approx (8|V_{cb}|G_F \Lambda_{BNV}^2)^{-2} \lesssim \mathcal{O}(5 \cdot 10^{-29})$$

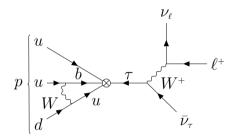
showing that direct observation is ruled out: $\bar{B} \rightarrow X \ell$

However...

what about operators with τ lepton? Not directly constrained by p decay!



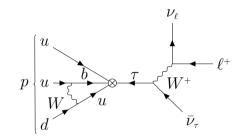
Both b and τ have to be virtual! Efficiently constrained by loop induced effects





Both b and au have to be virtual! Efficiently constrained by loop induced effects Estimate of this process $p \to \ell^+ \nu_\ell \bar{\nu}_\tau$ gives

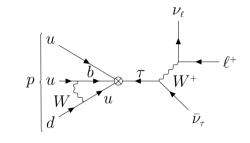
$$\Lambda_{\text{BNV}}\gtrsim (0.4\div 1.8)\cdot 10^6~\text{GeV}$$





Both b and au have to be virtual! Efficiently constrained by loop induced effects Estimate of this process $p \to \ell^+ \nu_\ell \bar{\nu}_\tau$ gives

$$\Lambda_{\text{BNV}} \gtrsim (0.4 \div 1.8) \cdot 10^6 \text{ GeV}$$





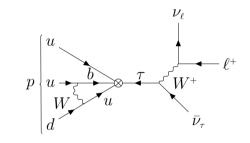
$$\mathcal{B}(\bar{B} \to X\tau) \lesssim (10^{-13} \div 10^{-15})$$

closer to detectability, but experimental efficiency in reconstructing au is much smaller...



Both b and au have to be virtual! Efficiently constrained by loop induced effects Estimate of this process $p \to \ell^+ \nu_\ell \bar{\nu}_\tau$ gives

$$\Lambda_{\text{BNV}} \gtrsim (0.4 \div 1.8) \cdot 10^6 \text{ GeV}$$





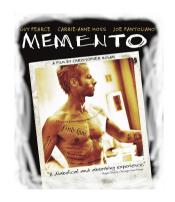
$$\mathcal{B}(\bar{B} \to X\tau) \lesssim (10^{-13} \div 10^{-15})$$

closer to detectability, but experimental efficiency in reconstructing au is much smaller...

$$\mathcal{B}(B^0 \to p\mu^-)|_{\text{exp}} < 2.6 \cdot 10^{-9}$$
 [LHCb 2022]



Left-Handed *b* **SMEFT Operators**





Left-Handed *b* **SMEFT Operators**

$$Q_{RR} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] [\widetilde{u}_s^c P_R \ell_t]$$

$$Q_{RL} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] [\widetilde{u}_s^c P_L \ell_t]$$

$$Q_{LR} = \varepsilon^{abc} [\widetilde{d}_p^a P_L u_r^b] \ [\widetilde{u}_s^c P_R \ell_t]$$

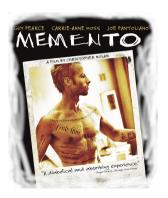
$$Q_{LL} = \varepsilon^{abc} [\widetilde{d}^a_p P_L u^b_r] \ [\widetilde{u}^c_s P_L \ell_t]$$

$$Q_{R\nu} = \varepsilon^{abc} [\tilde{d}_p^a P_R u_r^b] [\tilde{d}_s^c P_L \nu_t]$$

$$Q_{L\nu} = \varepsilon^{abc} [\widetilde{d}_p^a P_L u_r^b] [\widetilde{d}_s^c P_L \nu_t]$$

one loop SMEFT→WET matching known

[W. Dekens, P. Stoffer, 2019]





Left-Handed *b* **SMEFT Operators**

Wilson coefficients in the WET are not all independent, due to CKM rotation of the left-handed d quarks

$$Q_{RR} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] \ [\widetilde{u}_s^c P_R \ell_t]$$

$$Q_{RL} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] [\widetilde{u}_s^c P_L \ell_t]$$

$$Q_{LR} = \varepsilon^{abc} [\widetilde{d}^a_p P_L u^b_r] \ [\widetilde{u}^c_s P_R \ell_t]$$

$$Q_{LL} = \varepsilon^{abc} [\widetilde{d}_p^a P_L u_r^b] \ [\widetilde{u}_s^c P_L \ell_t]$$

$$Q_{R\nu} = \varepsilon^{abc} [\tilde{d}_p^a P_R u_r^b] [\tilde{d}_s^c P_L \nu_t]$$

$$Q_{L\nu} = \varepsilon^{abc} [\widetilde{d}_p^a P_L u_r^b] [\widetilde{d}_s^c P_L \nu_t]$$

one loop SMEFT→WET matching known

[W. Dekens, P. Stoffer, 2019]





Left-Handed *b* **SMEFT Operators**

Wilson coefficients in the WET are not all independent, due to CKM rotation of the left-handed *d* quarks

$$C_{RL}^{111} = C_{duq}^{111} \,,$$

$$C_{R\nu}^{111} = -V_{ud}C_{duq}^{111} - V_{cd}C_{duq}^{112} - V_{td}C_{duq}^{113} \,,$$

$$C_{R\nu}^{112} = -V_{us}C_{duq}^{111} - V_{cs}C_{duq}^{112} - V_{ts}C_{duq}^{113} ,$$

$$C_{R\nu}^{113} = -V_{ub}C_{duq}^{111} - V_{cb}C_{duq}^{112} - V_{tb}C_{duq}^{113} ,$$

$$Q_{RR} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] [\widetilde{u}_s^c P_R \ell_t]$$

$$Q_{RL} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] \ [\widetilde{u}_s^c P_L \ell_t]$$

$$Q_{LR} = \varepsilon^{abc} [\widetilde{d}^a_p P_L u^b_r] \ [\widetilde{u}^c_s P_R \ell_t]$$

$$Q_{LL} = \varepsilon^{abc} [\widetilde{d}_p^a P_L u_r^b] \ [\widetilde{u}_s^c P_L \ell_t]$$

$$Q_{R\nu} = \varepsilon^{abc} [\widetilde{d}_p^a P_R u_r^b] [\widetilde{d}_s^c P_L \nu_t]$$

$$Q_{L\nu} = \varepsilon^{abc} [\widetilde{d}_p^a P_L u_r^b] \ [\widetilde{d}_s^c P_L \nu_t]$$

one loop SMEFT→WET matching known

[W. Dekens, P. Stoffer, 2019]







Under the assumption that BNV occurs only if b quark is involved:

Scanned for the less constrained SMEFT operators

Under the assumption that BNV occurs only if b quark is involved:

- Scanned for the less constrained SMEFT operators
- Derived bounds for Λ_{BNV} from proton decay in three decay channels

Under the assumption that BNV occurs only if b quark is involved:

- Scanned for the less constrained SMEFT operators
- ullet Derived bounds for Λ_{BNV} from proton decay in three decay channels
- Showed $\mathcal{B}(\bar{B} \to X\ell) \lesssim \mathcal{O}(5 \cdot 10^{-29}) \Rightarrow$ undetectable!

Under the assumption that BNV occurs only if b quark is involved:

- Scanned for the less constrained SMEFT operators
- ullet Derived bounds for Λ_{BNV} from proton decay in three decay channels
- Showed $\mathcal{B}(\bar{B} o X\ell) \lesssim \mathcal{O}(5 \cdot 10^{-29}) \Rightarrow$ undetectable!
- For τ less restrictive $\mathcal{B}(\bar{B} \to X\tau) \lesssim 10^{-13} \div 10^{-15}$

Under the assumption that BNV occurs only if b quark is involved:

- Scanned for the less constrained SMEFT operators
- ullet Derived bounds for Λ_{BNV} from proton decay in three decay channels
- Showed $\mathcal{B}(\bar{B} o X\ell) \lesssim \mathcal{O}(5 \cdot 10^{-29}) \Rightarrow$ undetectable!
- For au less restrictive $\mathcal{B}(\bar{B} \to X au) \lesssim 10^{-13} \div 10^{-15}$

Take-home message:

Direct detection of BNV *B* decays is impossible!

Under the assumption that BNV occurs only if b quark is involved:

- Scanned for the less constrained SMEFT operators
- \bullet Derived bounds for Λ_{BNV} from proton decay in three decay channels
- Showed $\mathcal{B}(\bar{B} o X\ell) \lesssim \mathcal{O}(5 \cdot 10^{-29}) \Rightarrow$ undetectable!
- For τ less restrictive $\mathcal{B}(\bar{B} \to X\tau) \lesssim 10^{-13} \div 10^{-15}$

Take-home message:

Direct detection of BNV *B* decays is impossible!

me: "... we should be careful, only in the assumption of dim-6 interactions..."

Under the assumption that BNV occurs only if b quark is involved:

- Scanned for the less constrained SMEFT operators
- \bullet Derived bounds for Λ_{BNV} from proton decay in three decay channels
- Showed $\mathcal{B}(\bar{B} \to X\ell) \lesssim \mathcal{O}(5 \cdot 10^{-29}) \Rightarrow$ undetectable!
- For au less restrictive $\mathcal{B}(\bar{B} \to X au) \lesssim 10^{-13} \div 10^{-15}$

Take-home message:

Direct detection of BNV *B* decays is impossible!

me: "... we should be careful, only in the assumption of dim-6 interactions..."

MB: "you become even more famous if you state wrong no-go theorems!"

Thank You!



\dots in B Physics

Theoretical predictions are affected by the **non-perturbative** nature of hadronic QCD

However nature provided an intrinsic perturbative scale $m_b \sim 5 \text{ GeV}$



 \Rightarrow B decays employing **Effective Field Theories** (HQET, SCET, ...) to separate perturbative physics from universal non-perturbative inputs

